

Physics 6320
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Simple Harmonic Oscillator
Part IV: Sample Problems

1 Review.

Last time we derived wavefunctions for the first few SHO states:

$$\begin{aligned}\phi_0(x) = \langle x | 0 \rangle &= \sqrt{\frac{1}{x_0\sqrt{\pi}}} e^{-x^2/2x_0^2} \\ \phi_1(x) = \langle x | 1 \rangle &= \frac{\sqrt{2}}{x_0} x\phi_0(x) \\ \phi_2(x) = \langle x | 2 \rangle &= \frac{1}{\sqrt{2}x_0^2}(2x^2 - x_0^2)\phi_0(x) \\ &\dots \\ &\dots\end{aligned}\tag{1}$$

We can verify that these satisfy

$$\int_{-\infty}^{\infty} \phi_n(x)^* \phi_m(x) dx = \delta_{mn}\tag{2}$$

as they must, by using the integral formula

$$\int_{-\infty}^{\infty} x^{2n} e^{-x^2} dx = \Gamma(n + 1/2)$$

and the gamma function properties

$$\Gamma(n + 1) = n\Gamma(n) \quad \text{and} \quad \Gamma(1/2) = \sqrt{\pi}.$$

2 Parity of oscillator wavefunctions.

Since the SHO potential $V = \frac{1}{2}\omega^2 x^2$ is invariant under the transformation $x \rightarrow -x$, all the energy eigenstates $\phi_n(x)$ have definite parity. In fact, $\phi_n(-x) = (-1)^n \phi_n(x)$. Formally we say that the parity operator commutes with the hamiltonian, and therefore the energy and the parity are compatible: we can have simultaneous eigenstates. This does not mean that we cannot have states of mixed parity, but they will not be energy eigenstates. We consider such an example next.

3 Example 1.

Suppose at time $t = 0$ we prepare the oscillator in state $|\psi(0)\rangle$ described by the wavefunction

$$\langle x | \psi(0) \rangle = \psi(x, 0) = C(1 + x/x_0) e^{-x^2/2x_0^2}.\tag{3}$$

How will this state develop with time?

First let's get the normalization, using $\xi = x/x_0$ as usual:

$$\begin{aligned}\int_{-\infty}^{\infty} |\psi(x, 0)|^2 dx &= C^2 x_0 \int_{-\infty}^{\infty} (1 + \xi)^2 e^{-\xi^2} d\xi = C^2 x_0 \{\Gamma(1/2) + 0 + \Gamma(3/2)\} \\ &= \frac{3}{2} \sqrt{\pi} x_0 C^2 = 1. \\ \text{So } C^2 &= \frac{2}{3x_0\sqrt{\pi}}.\end{aligned}$$

Now let's try to express our given $\psi(x, 0)$ in terms of energy eigenfunctions so we can get the time dependence. It looks like a linear combination of ϕ_0 and ϕ_1 . Let's try that: from (1) we have

$$(1 + \xi)\phi_0 = \phi_0 + \sqrt{\frac{1}{2}}\phi_1. \quad (4)$$

But we are given

$$\psi(x, 0) = C(1 + \xi)e^{-\xi^2/2} = C\sqrt{x_0\sqrt{\pi}}(1 + \xi)\phi_0 \quad (5)$$

So

$$\begin{aligned} \psi(x, 0) &= C\sqrt{x_0\sqrt{\pi}}(1 + \xi)\phi_0 = \sqrt{\frac{2}{3x_0\sqrt{\pi}}}\sqrt{x_0\sqrt{\pi}}(1 + \xi)\phi_0 \\ &= \sqrt{\frac{2}{3}}(1 + \xi)\phi_0 = \sqrt{\frac{2}{3}}\phi_0 + \sqrt{\frac{1}{3}}\phi_1. \end{aligned} \quad (6)$$

So now the time dependence is obvious,

$$\psi(x, t) = \sqrt{\frac{2}{3}}\phi_0(x)e^{-i\omega_0 t} + \sqrt{\frac{1}{3}}\phi_1(x)e^{-i\omega_1 t} \quad (7)$$

Here of course the eigenfrequencies are

$$\omega_n = \frac{E_n}{\hbar} = (n + 1/2)\omega \quad \text{where} \quad \omega = \sqrt{k/m} = \text{classical frequency} \quad (8)$$

Where will we find this particle? Of course when the system is in an energy eigenstate, $|\phi(-x)|^2 = |\phi(x)|^2$ so that $\langle \text{phi}_n | x | \phi_n \rangle = 0$. But our state $\psi(x, t)$ is neither even nor odd, it is a mixture, so $\langle x \rangle$ need not be zero. Let's calculate it:

$$\begin{aligned} \psi(x, t) &= \sqrt{\frac{2}{3}}\phi_0(x)e^{-i\omega t/2} + \sqrt{\frac{1}{3}}\phi_1(x)e^{-3i\omega t/2} \\ &= \sqrt{\frac{2}{3}}e^{-i\omega t/2} (1 + \xi e^{-i\omega t}) \phi_0(x) \\ \langle \psi | x | \psi \rangle &= \int_{-\infty}^{\infty} \psi^*(x, t) x \psi(x, t) dx \\ &= \frac{2}{3}x_0^2 \int_{-\infty}^{\infty} |1 + \xi e^{-i\omega t}|^2 |\phi_0(x)|^2 \xi d\xi \\ &= \frac{4}{3}x_0^2 \cos(\omega t) \int_{-\infty}^{\infty} |\phi_0(x)|^2 \xi^2 d\xi \\ &= \frac{4x_0}{3\sqrt{\pi}} \cos(\omega t) \int_{-\infty}^{\infty} \xi^2 e^{-\xi^2} d\xi = \frac{2}{3}x_0 \cos(\omega t) \end{aligned}$$

So in this state, the mean value of the particle's position oscillates about zero, with the classical frequency, and with amplitude $2x_0/3$. This is an example of Ehrenfest's theorem, which says that expectation values follow classical equations of motion. In fact, no matter what initial state we prepare, if it has a non-zero value of $\langle x \rangle$, then that value will oscillate with the classical frequency.

4 Oscillation of a gaussian packet.

Encouraged by our previous example, let us try to prepare our oscillator in a minimum-uncertainty state, sharply peaked about some point $x = y$. That is, we pick

$$\psi(x, 0) = \sqrt{\frac{1}{w\sqrt{\pi}}} e^{-(x-y)^2/2w^2} \quad (9)$$

As we already know, this gives a normalized probability function, peaked at y , with width w .

Now can we expand this in terms of oscillator eigenfunctions?

It turns out to be quite messy for an arbitrarily chosen value of the width w . But it works out very nicely if we pick $w = x_0$. In other words, suppose we start with the ground state wavefunction ϕ_0 , but displaced to be centered around $x = y$ instead of $x = 0$.

We want

$$c_n = \langle \phi_n | \psi(0) \rangle \quad \text{so that} \quad |\psi(t)\rangle = \sum_n c_n e^{-i\omega_n t} |\phi_n\rangle \quad (10)$$

So we need the integral

$$c_n = \langle \phi_n | \psi(0) \rangle = \sqrt{\frac{1}{w\sqrt{\pi}}} \int_{-\infty}^{\infty} \phi_n(x) e^{-(x-y)^2/2a_0^2} dx \quad (11)$$

This can be done readily if we take advantage of some facts about Hermite polynomials. First of course there is the orthonormality relation,

$$\int_{-\infty}^{\infty} H_n(x) H_m(x) e^{-x^2} dx = \delta_{nm} \sqrt{\pi} n! 2^n. \quad (12)$$

(Note that there is a typo in Sakurai (A.4.5) which gives $2^n!$ instead of $2^n \cdot n!$.)

There is a very useful function called the **generating function** for Hermite polynomials. (We will see this kind of thing in other cases, such as Legendre polynomials.) The generating function for Hermite polynomials is

$$F(x, s) = e^{-s^2 + 2xs} = \sum_{n=0}^{\infty} H_n(x) \frac{s^n}{n!}. \quad (13)$$

In fact we can **define** the Hermite polynomials as the coefficients in the power series of $F(x, s)$. From this we can get two useful formulas. First, consider the integral

$$\int_{-\infty}^{\infty} F(x, s) e^{-x^2} H_n(x) dx = \int_{-\infty}^{\infty} e^{-(x-s)^2} H_n(x) dx = \sum_{m=0}^{\infty} \frac{s^m}{m!} \int_{-\infty}^{\infty} H_n(x) H_m(x) e^{-x^2} dx \quad (14)$$

Now use of (12) gives

$$\int_{-\infty}^{\infty} e^{-(x-s)^2} H_n(x) dx = \sqrt{\pi} (2s)^n \quad (15)$$

This result allows us to do the integral in (11) to get

$$c_n = \langle \phi_n | \psi(0) \rangle = \sqrt{\frac{1}{n! 2^n}} (z/x_0)^n e^{-z^2/4x_0^2} \quad (16)$$

Note this means our initial function has amplitudes for infinitely large energies. Now we can use the generating function **again** to do the sum in (10)! The result is a little bit messy, but to get the time-dependent probability distribution we need its modulus squared, which turns out to be just

$$|\psi(x, t)|^2 = \frac{1}{x_0 \sqrt{\pi}} e^{-(x - z \cos \omega t)^2 / x_0^2} \quad (17)$$

In other words we have reconstructed the ground-state function $\phi_0(x)$, but now as a function of $x(t) = z \cos(\omega t)$, where again ω is the classical frequency.