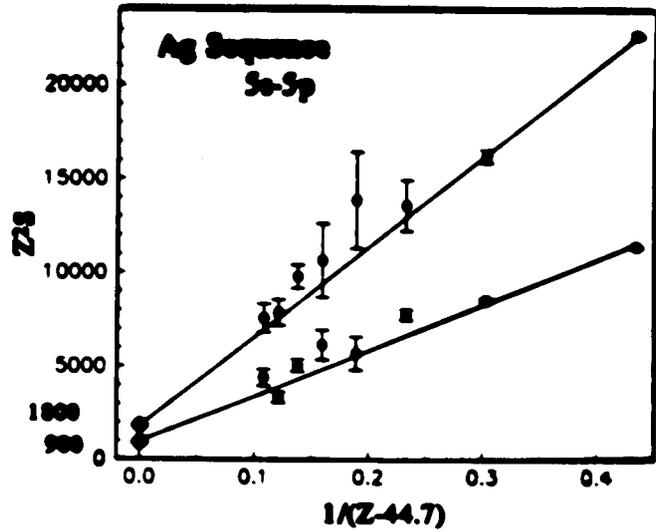
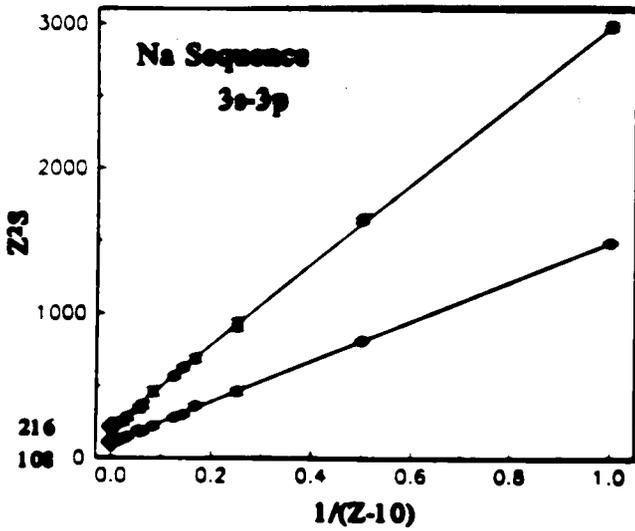
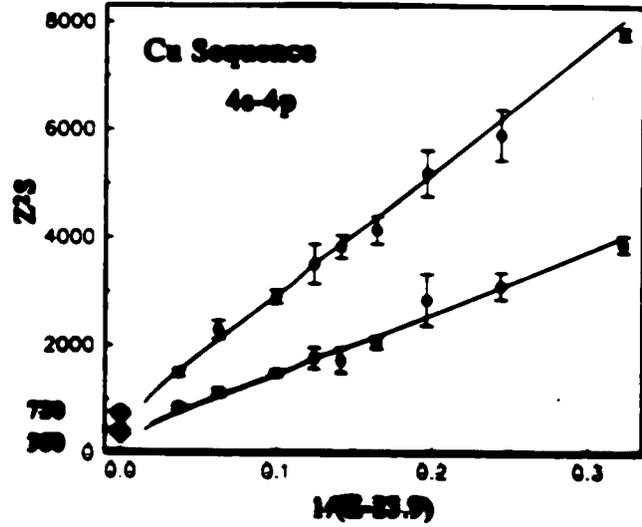
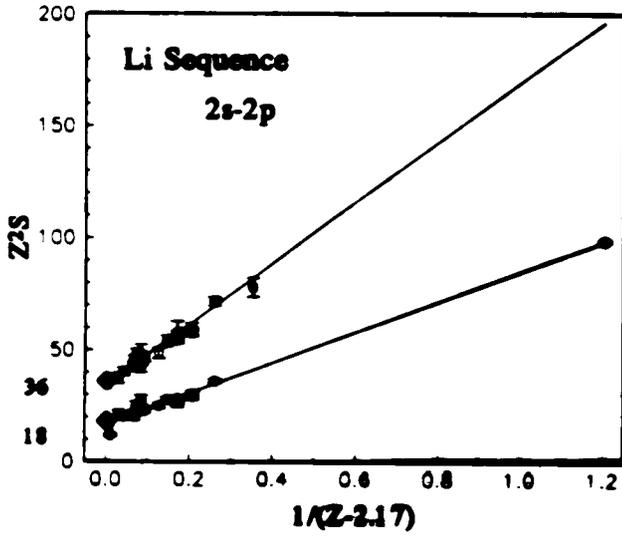


$$Z^2 S = S_H$$

H-like

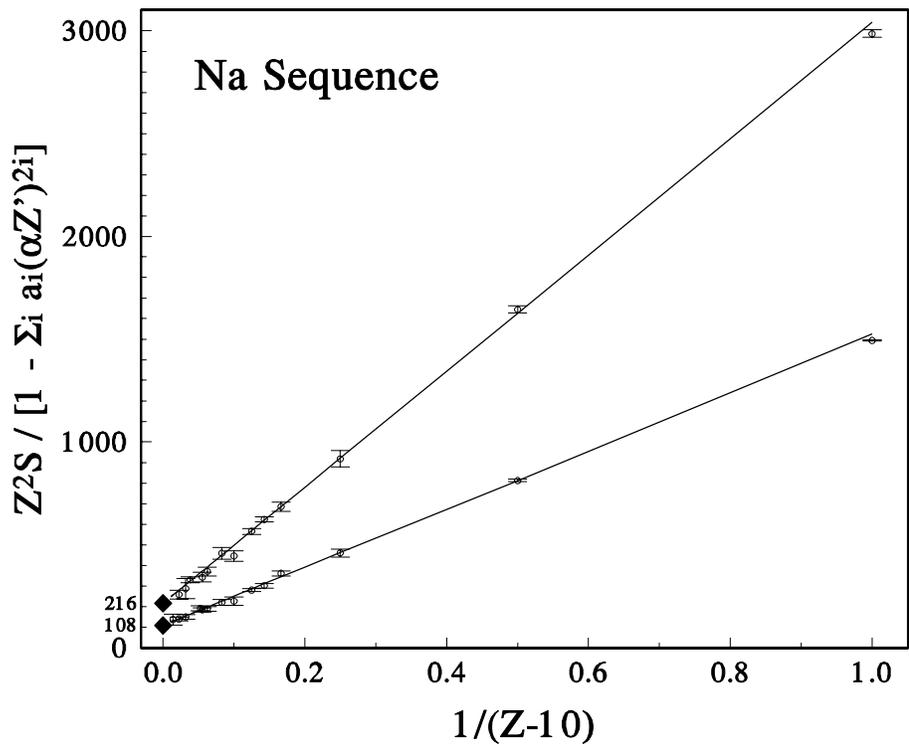
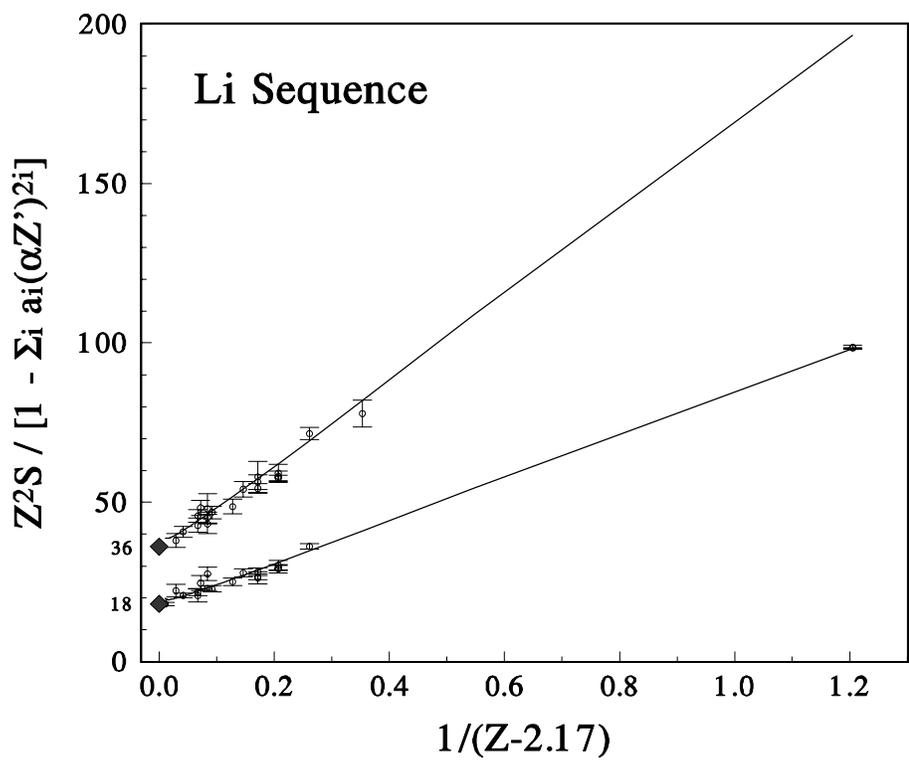
$$\longrightarrow S_H + \frac{b}{Z-C}$$

Alkalilike



$$\blacklozenge S_H(ms-np) = \frac{3}{4} n^2 (n^2 - 1) (2J + 1)$$

$$S_{ik} = \left[\frac{\lambda(\text{\AA})}{1265.38} \right]^3 \frac{g_i B_{ik}}{\tau_i(\text{ns})}$$



--- Johnson, Liu & Sapirstein, *Atom.Data Nucl.Data Tables* 64, 279-300 (1996).

Relativistic Hydrogenlike Line Strengths

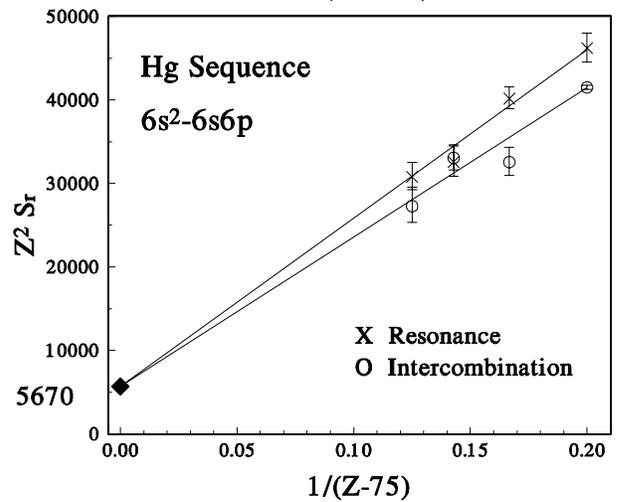
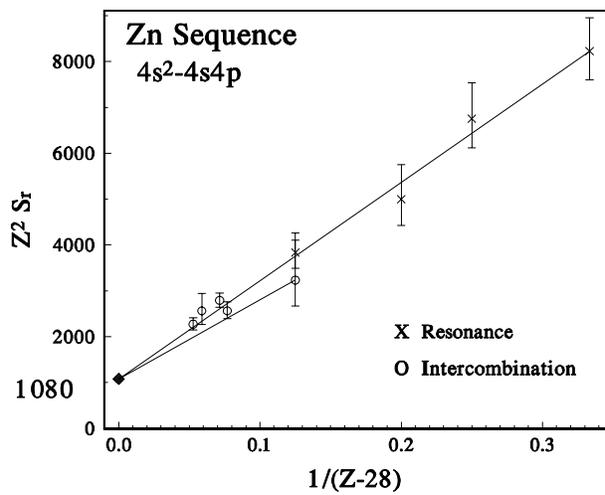
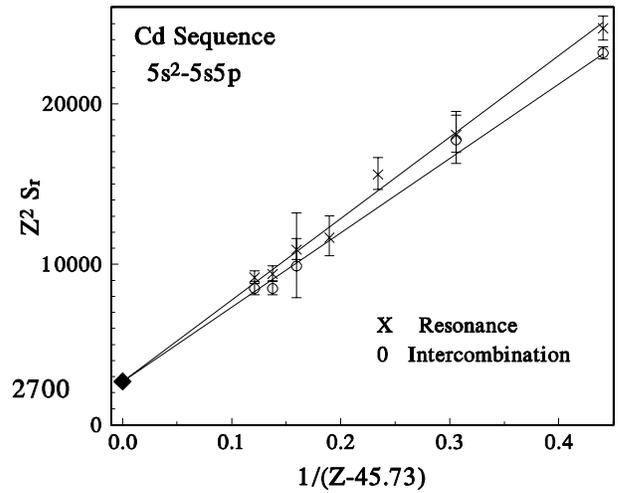
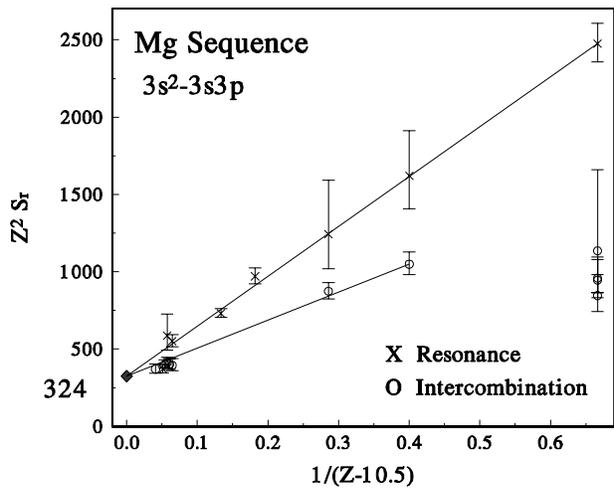
$$Z^2 S_H^R(Z) = Z^2 S_H \left[1 - \sum_{i=1}^{\infty} a_i (\alpha Z)^{2i} \right]$$

TABLE I. Constants for relativistic hydrogenlike line strengths.

Transition	$S_H(1)$	a_1	a_2	a_3	a_4	a_5
$2s_{1/2} - 2p_{1/2}$	18	5/6	-1/48	1/96	7/768	11/1536
$2s_{1/2} - 2p_{3/2}$	36	1/3	0.110187	0.059476	0.037032	0.024925
$3s_{1/2} - 3p_{1/2}$	108	7/12	5/144	7/288	37/2304	53/4608
$3s_{1/2} - 3p_{3/2}$	216	19/72	0.139267	0.082219	0.050840	0.033291
$4s_{1/2} - 4p_{1/2}$	360	9/20	3/64	17/640	87/5120	123/10240
$4s_{1/2} - 4p_{3/2}$	720	103/480	0.147237	0.093466	0.058853	0.038459
$5s_{1/2} - 5p_{1/2}$	900	11/30	29/600	31/1200	157/9600	221/19200
$5s_{1/2} - 5p_{3/2}$	1800	9/50	0.150012	0.100133	0.064068	0.041970

New plotting parameter

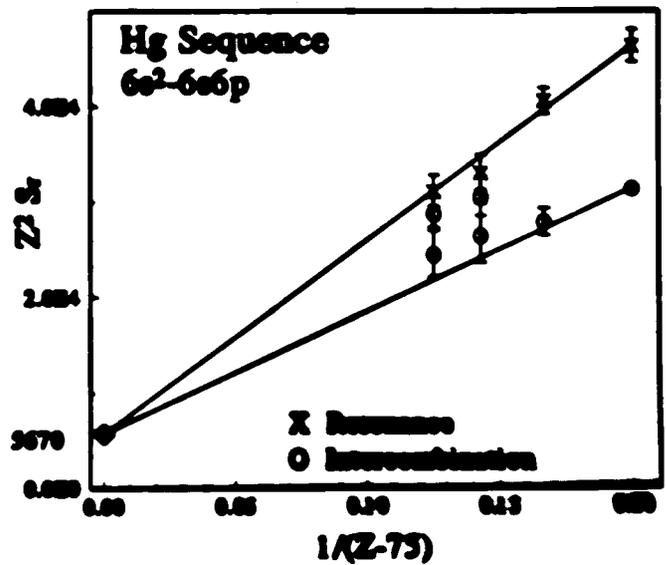
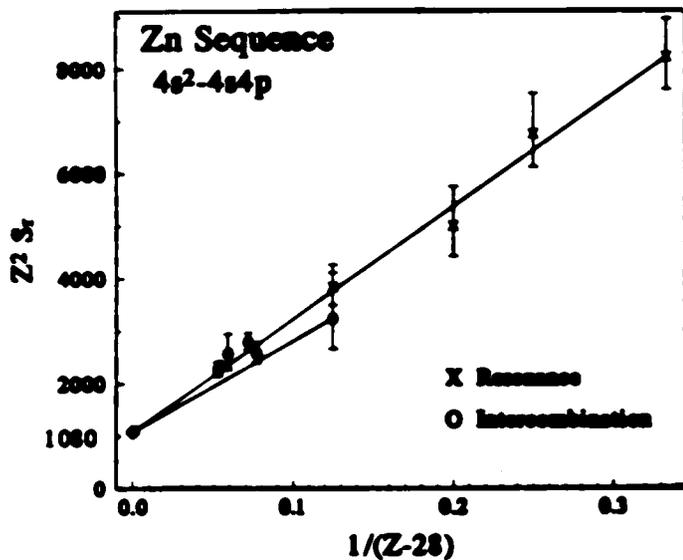
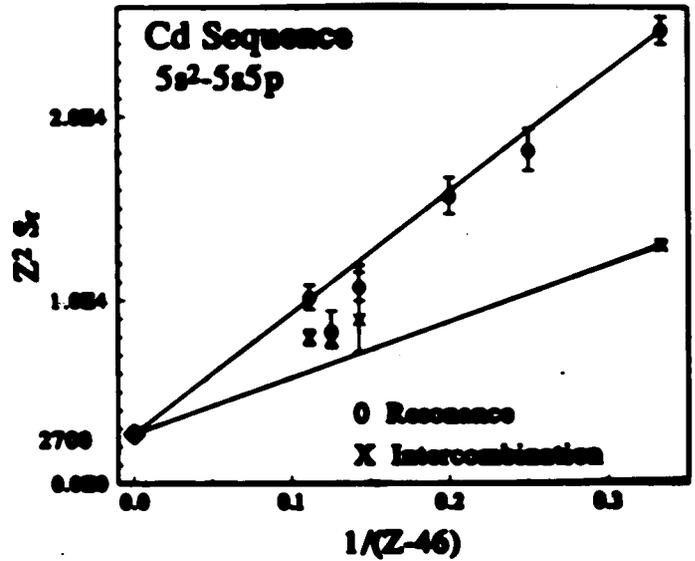
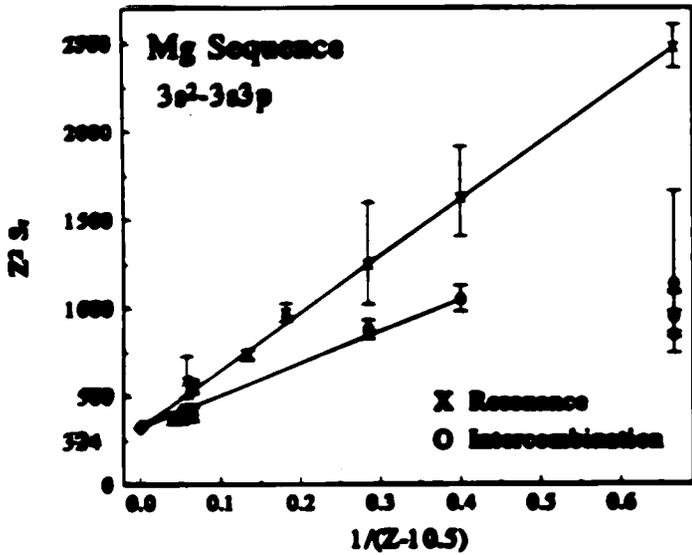
$$\frac{Z^2 S}{1 - \sum_i a_i [\alpha(Z-C)]^{2i}} \cong Z^2 S_H \left[A' + \frac{B'}{(Z-C)} \right]$$



$$S_r(\text{Res}) = \frac{S(\text{Res})}{\cos^2 \theta} ;$$

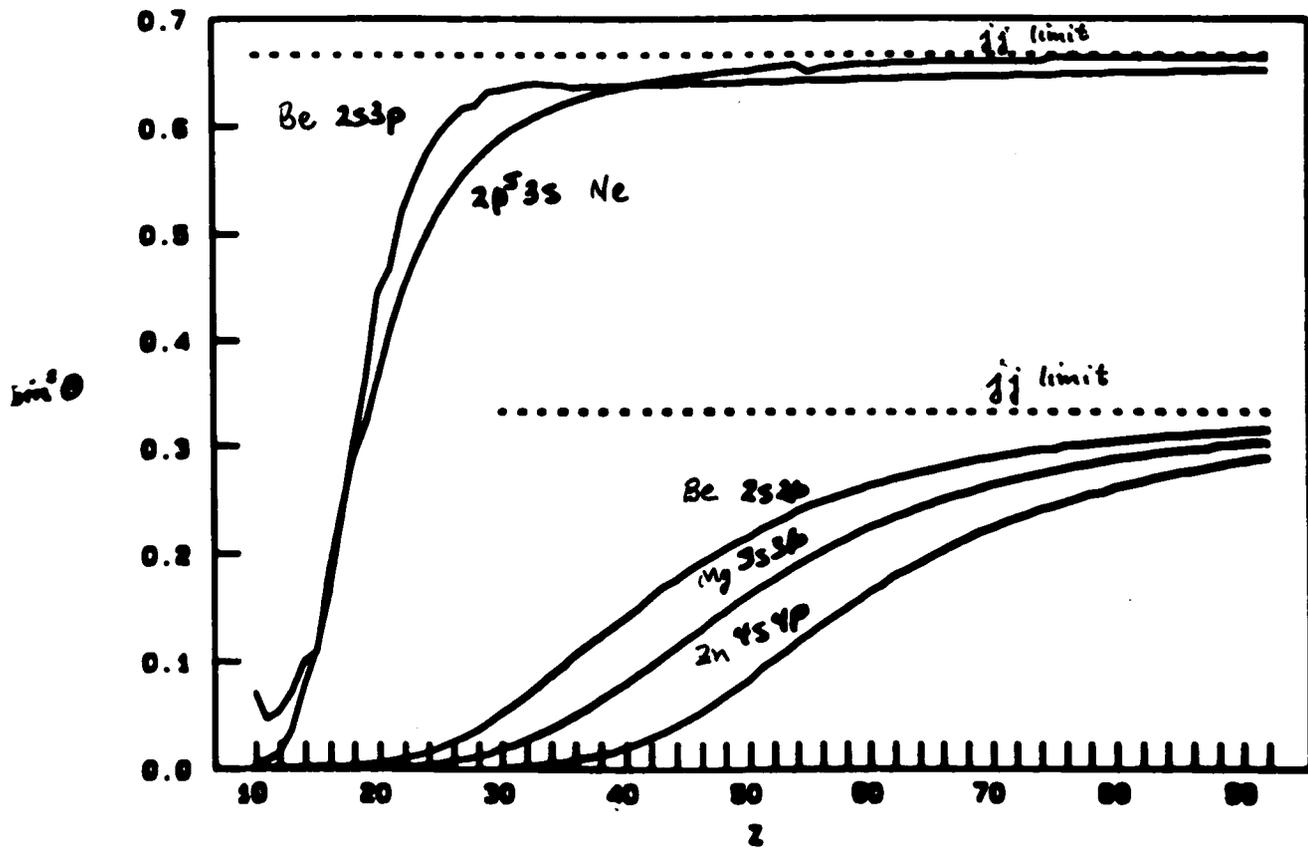
$$S_r(\text{Int}) = \frac{S(\text{Int})}{\sin^2 \theta}$$

θ determined from energy level data

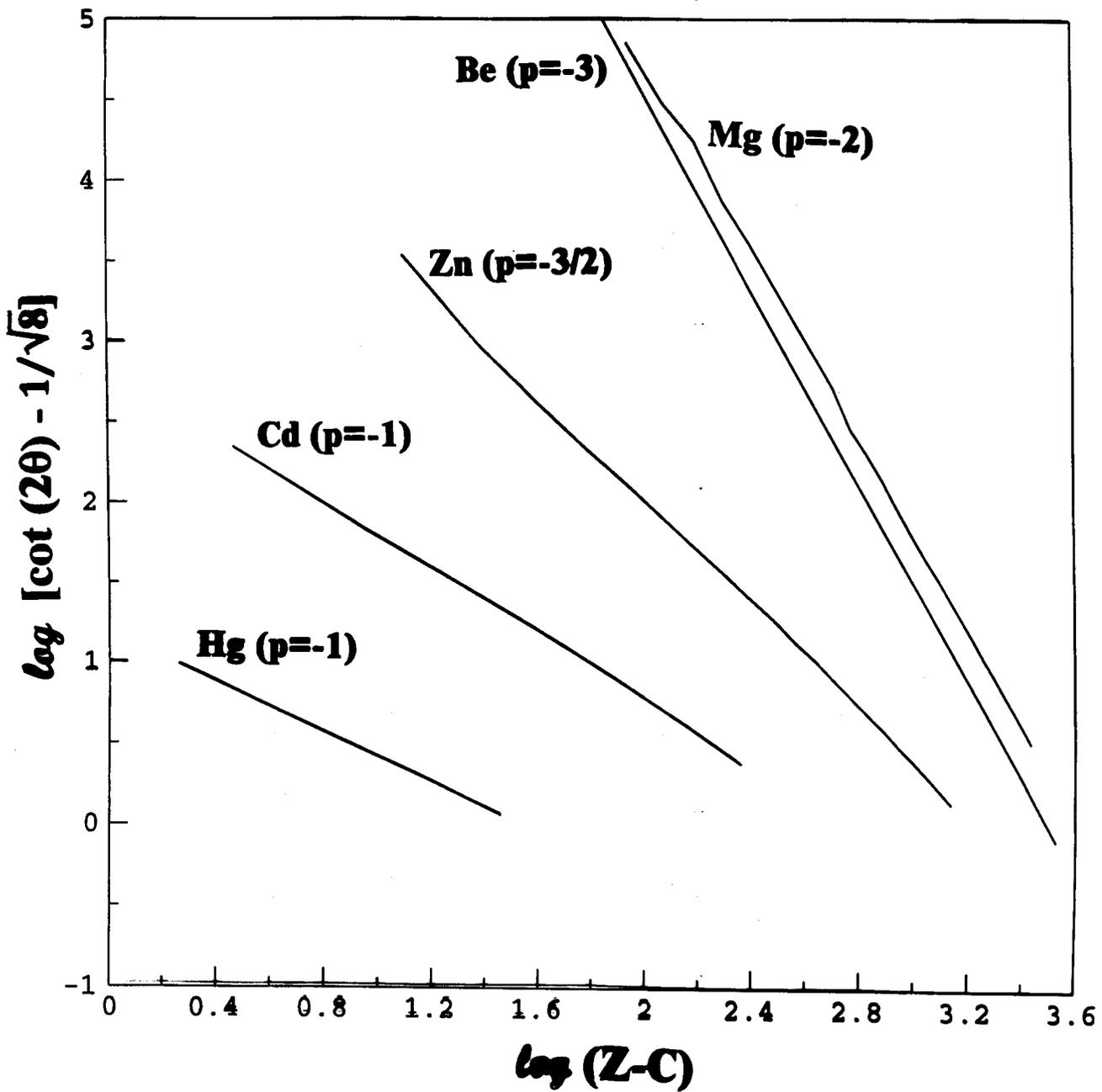


$$S_H = \frac{9}{2} n^2 (n^2 - 1)$$

Singlet-Triplet Mixing Angles



$$\cot(2\theta) = 1/\sqrt{8} + A(Z-C)^p$$



ns np

Hugh Wolfe, PR 41, 443 (1932)

3P_1 G.W. King 3P_2 J.H. Van Vleck 1P_1 PR 56, 464 (1939)

	3P_0	3P_1	3P_2	1P_1
3P_0	$E_0 - G_1 - \mu_1$	0	0	0
3P_1	0	$E_0 - G_1 - \mu_1/2$	0	$\mu_2/\sqrt{2}$
3P_2	0	0	$E_0 - G_1 + \mu_1/2$	0
1P_1	0	$\mu_2/\sqrt{2}$	0	$E_0 + G_1$

This is diagonal in the representation

$$|^3P_0'\rangle = |^3P_0\rangle$$

$$|^3P_1'\rangle = \cos\theta_1 |^3P_1\rangle - \sin\theta_1 |^1P_1\rangle$$

$$|^3P_2'\rangle = |^3P_2\rangle$$

$$|^1P_1'\rangle = \sin\theta_1 |^3P_1\rangle + \cos\theta_1 |^1P_1\rangle$$

with energies

$$|^3P_0'\rangle = E_0 - G_1 - \mu_1$$

$$|^3P_1'\rangle = E_0 - \mu_1/4 - \Delta$$

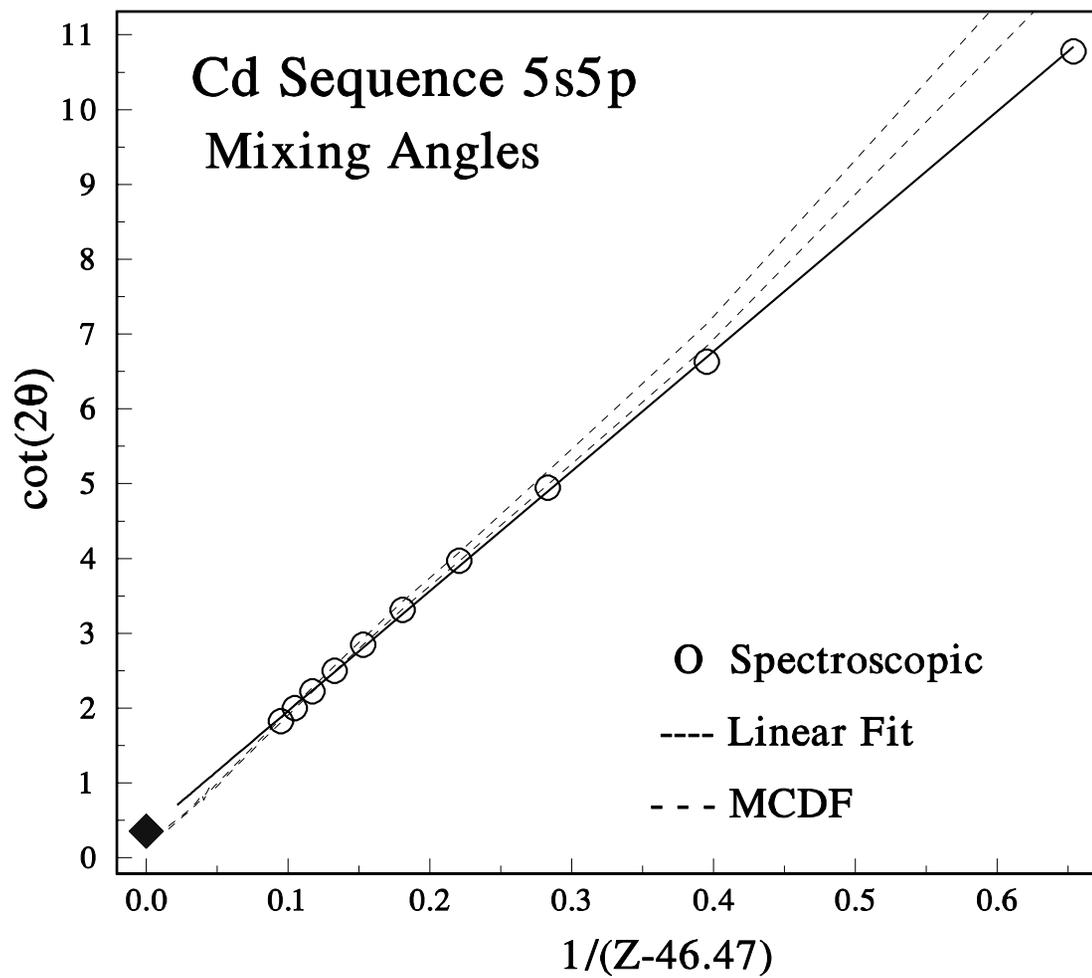
$$|^3P_2'\rangle = E_0 - G_1 + \mu_1/2$$

$$|^1P_1'\rangle = E_0 - \mu_1/4 + \Delta$$

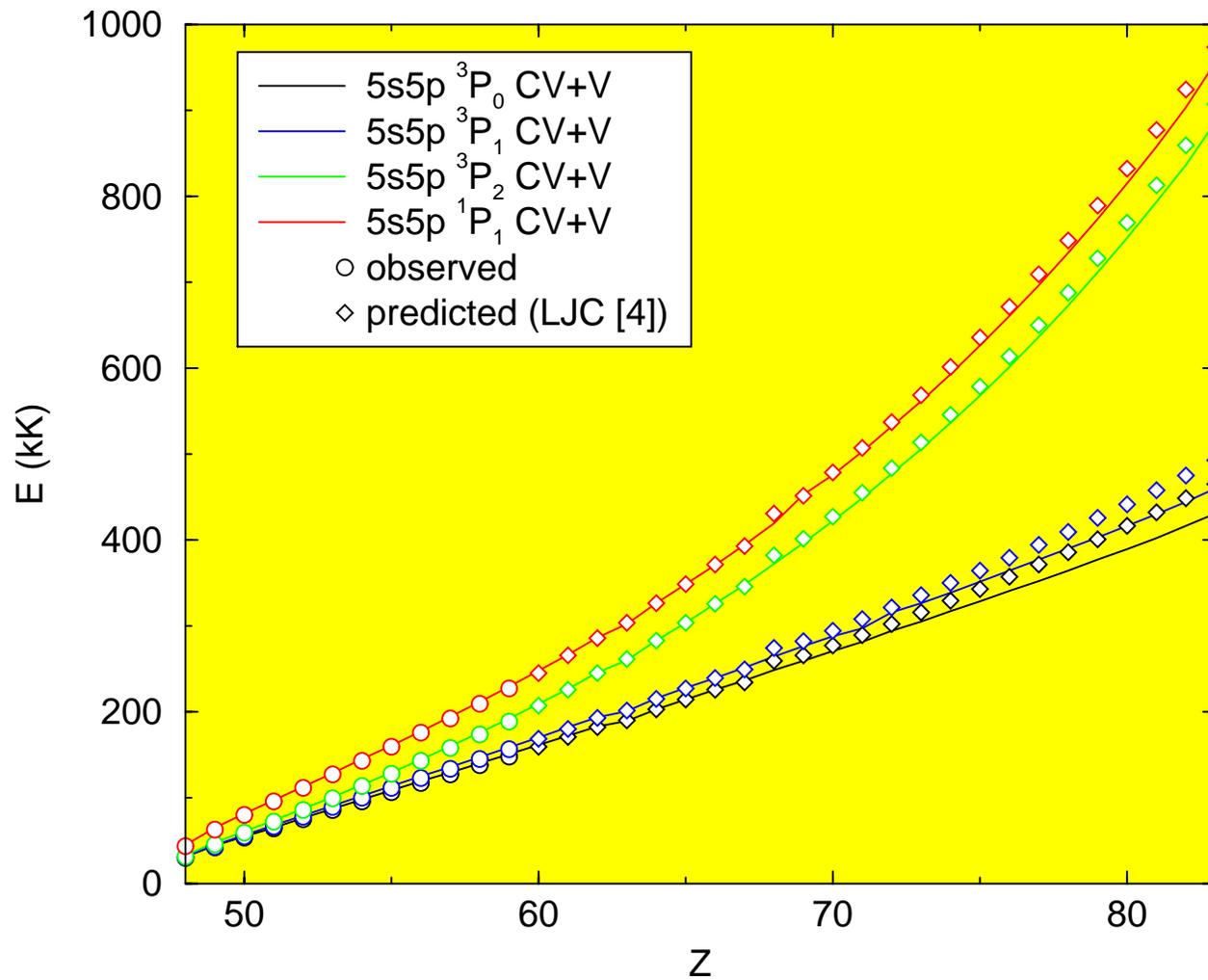
$$\Delta \equiv \sqrt{(G_1 + \mu_1/4)^2 + \mu_2^2/2}$$

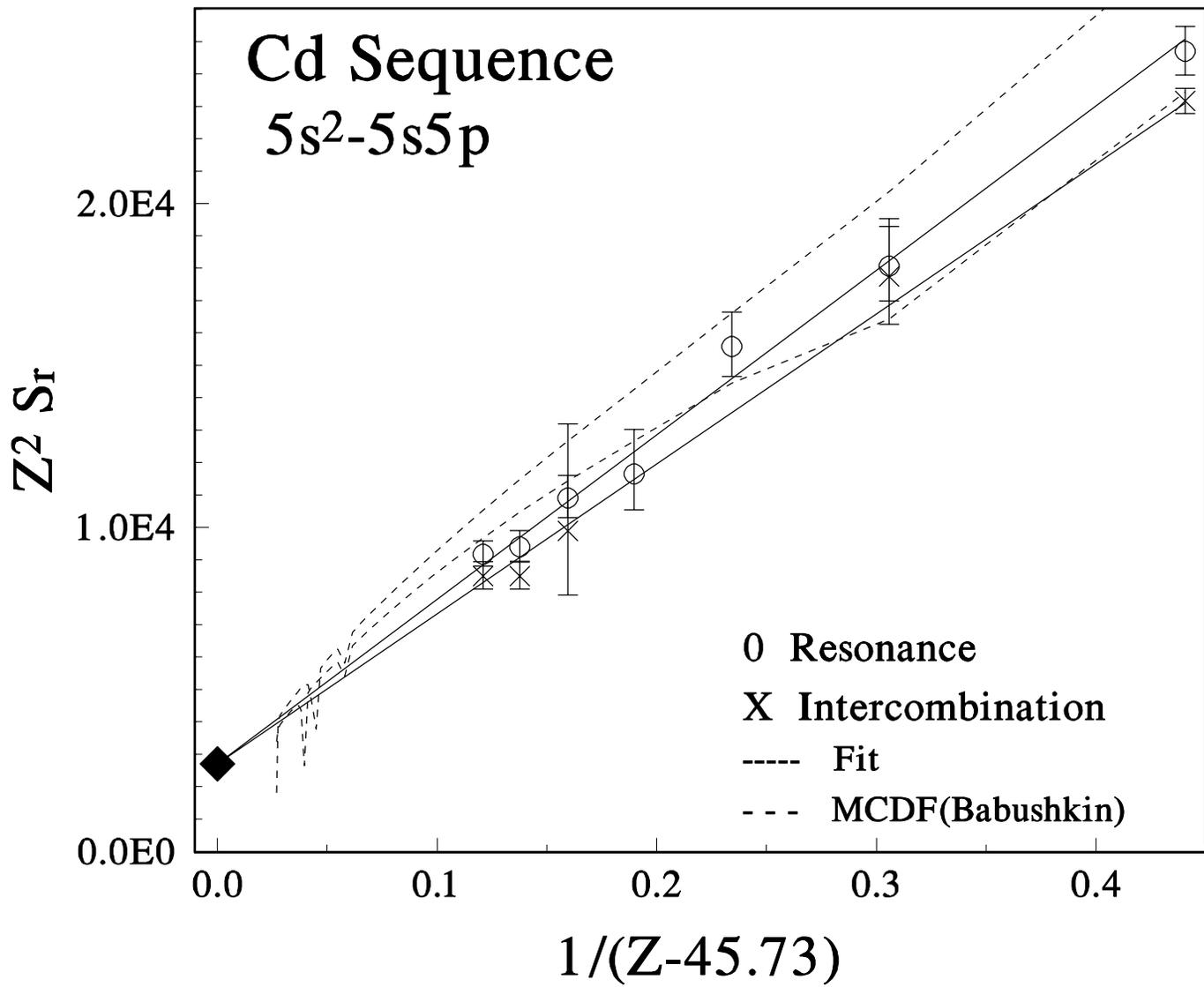
The determine the mixing angle

$$\cot 2\theta = \frac{4G_1 + \mu_1}{\sqrt{2}\mu_2}$$

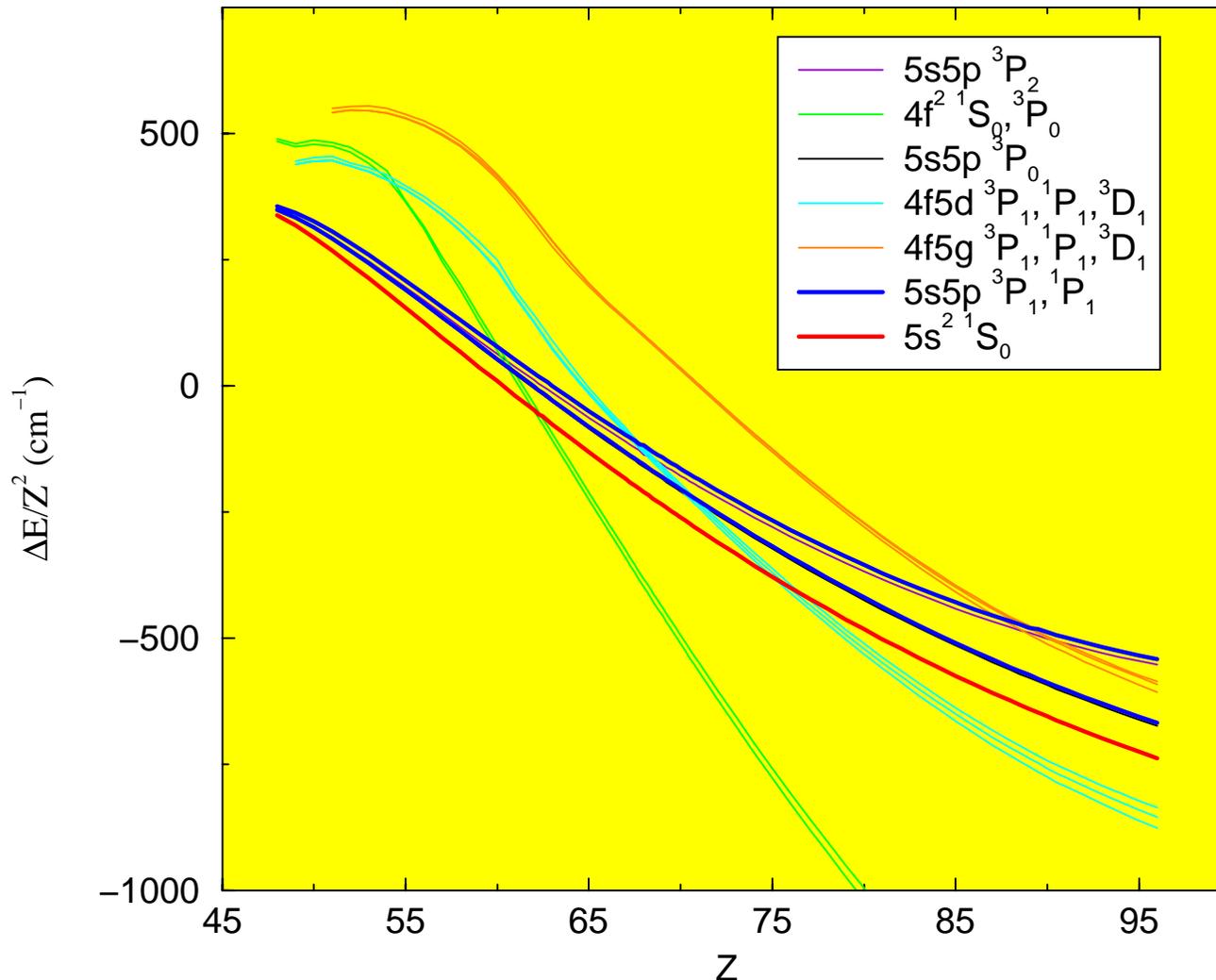


$5s^2\ ^1S_0 - 5s5p\ ^{1,3}P_1$ energy intervals





Energies relative to $4d^{10}5s^2S_{1/2}$ valence correlation only



Many more cross-ings occur in the CV+V calculation at high Z, as the **f** correlation orbital becomes the spectroscopic **4f** orbital in CSF's such as



TABLE III. Results for the intercombination transition $5s^2\ ^1S_0-5s5p\ ^3P_1$. Parentheses indicate quoted uncertainties in the experimental values as propagated from Table I.

Z ion	λ (Å)	A_{ik} (ns ⁻¹)					
		Expt.	MCDHF ^a		MCDHF ^b		SE ^c
			Coulomb	Babushkin	Coulomb	Babushkin	
48 Cd	3262	0.000418(7)	0.000348	0.000343	0.000449	0.000356	0.000419
49 In	2307	0.0023(2)	0.00212	0.00197	0.00244	0.00210	0.00210
50 Sn	1812		0.00676	0.00606	0.00701	0.00649	0.00614
51 Sb	1499		0.0154	0.0142	0.0158	0.0147	0.0137
52 Te	1282	0.025(6)	0.0304	0.0279	0.0310	0.0290	0.0261
53 I	1120	0.041(2)	0.0535	0.0490	0.0539	0.0509	0.0445
54 Xe	966	0.071(4)	0.0862	0.0793	0.0868	0.0822	0.0702
55 Cs	896		0.132	0.120	0.131	0.125	0.104
56 Ba	814		0.189	0.175	0.189	0.180	0.148
57 La	746		0.256	0.256	0.263	0.251	0.202
58 Ce	688				0.352	0.337	0.268
59 Pr	637				0.460	0.441	0.346
60 Nd	593				0.586	0.563	0.439
61 Pm	554				0.730	0.705	0.547
62 Sm	518				0.866	0.860	0.677
63 Eu	494				0.789	0.869	0.781
64 Gd	465				1.17	1.21	0.950
65 Tb	439				1.45	1.45	1.12
66 Dy	417				1.73	1.70	1.31
67 Ho	401				2.04	1.96	1.47

^aBiémont *et al.* [15].^bThis work, theoretical.^cThis work, semiempirical, from fits shown in Figs. 3 and 4 and Eqs. (21) and (22).

sition data could be effectively linearized by the same value of C . Therefore the fits, subject to constraints $S_0=2700$ and $C(\text{Res})=C(\text{Int})$, yielded $C=45.73$, $B(\text{Res})=50709$, and $B(\text{Int})=46274$. Summarizing, the predicted line strengths can be specified by

$$S(\text{Res}) = \left(2700 + \frac{50709}{Z-45.73} \right) \left(\frac{\cos \theta}{Z} \right)^2, \quad (20)$$

$$S(\text{Int}) = \left(2700 + \frac{46274}{Z-45.73} \right) \left(\frac{\sin \theta}{Z} \right)^2, \quad (21)$$

$$\cot 2\theta = 0.298 + 16.628/(Z-46.41). \quad (22)$$

The fitting constants can subsequently be sharpened as additional lifetime and energy level measurements become available. However it is significant to note that Eqs. (20)–(22) summarize, in very economical form, all of the information that is presently known concerning the line strengths of these Cd-like transitions.

The resonance and intercombination transition probability rates predicted by this semiempirical linearization are tabulated in Tables II and III for $48 \leq Z \leq 67$, together with the wavelengths and the MCDHF calculations. For $Z > 50$, the

results of the present calculation agree with the predictions of the considerably more elaborate theoretical model of Ref. [15]. This agreement indicates that the present approach should be sufficient to characterize the isoelectronic trends predicted by the MCDHF method at intermediate values of Z ; that is, away from the neutral- Z region that is complicated by extensive electron correlation effects, not yet in the high- Z region that is complicated by configuration interaction resulting from level crossings due to the contraction of the $4f$ shell. Although the quantitative description of the high- Z region would require a considerably more thorough theoretical effort, sharp deviations of our theoretical results from the smooth semiempirical isoelectronic trend in the high- Z region indicate the possible breakdown of the semiempirical formulation based on the single-configuration approach.

VI. CONCLUSIONS

The approximate linearities observed in the measured data through their expositions as $\cot 2\theta$ and Z^2S , vs $1/(Z-C)$ (with C optimally chosen) have been verified for the Cd sequence by MCDHF calculations in the region $48 \leq Z \leq 67$. This confirms the suggestion [4] that a few accurate lifetime measurements for the $5s5p\ ^1P_1$ and $5s5p\ ^3P_1$ levels can be

Contrast Cd and Mg Isoelectronic Sequences

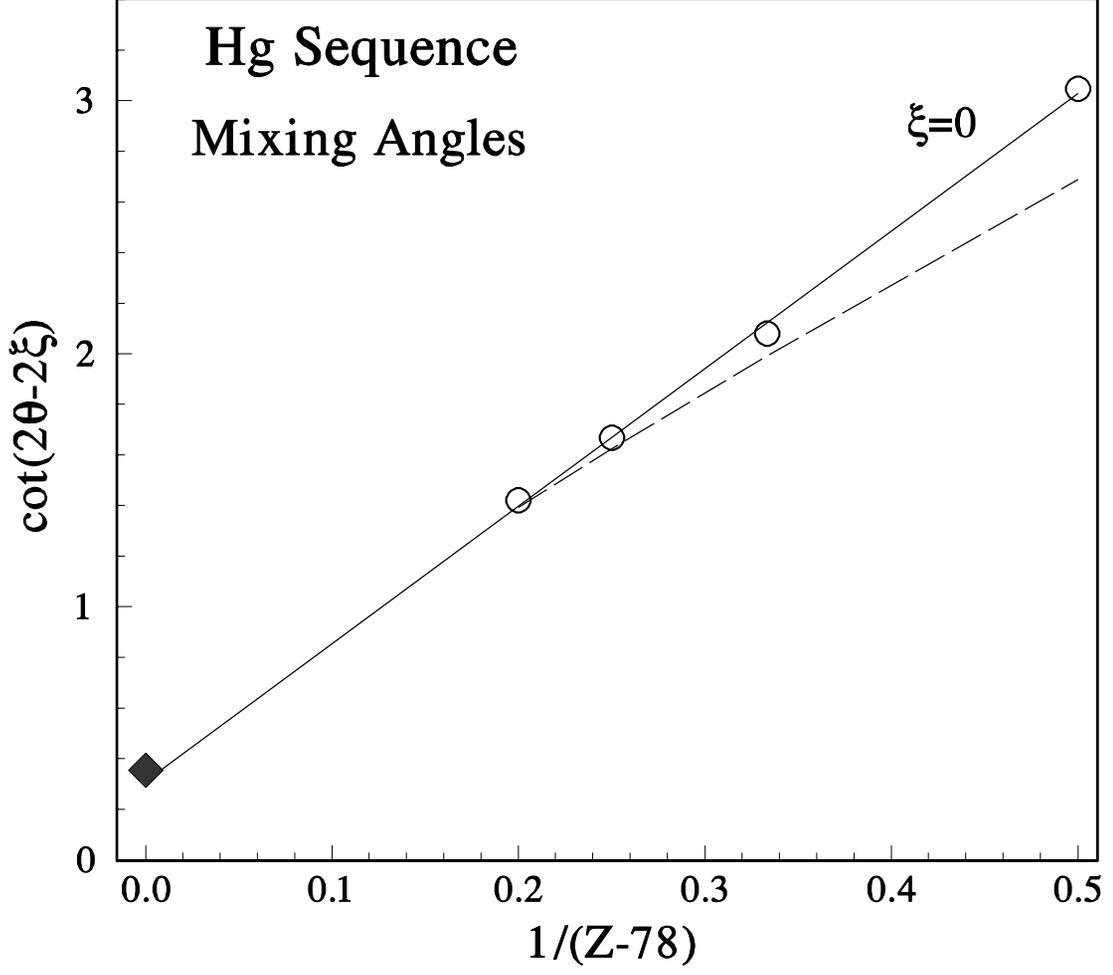
Cd: *48 electrons*
theoretically challenging
midrange Z, many stable isoelectronic ions

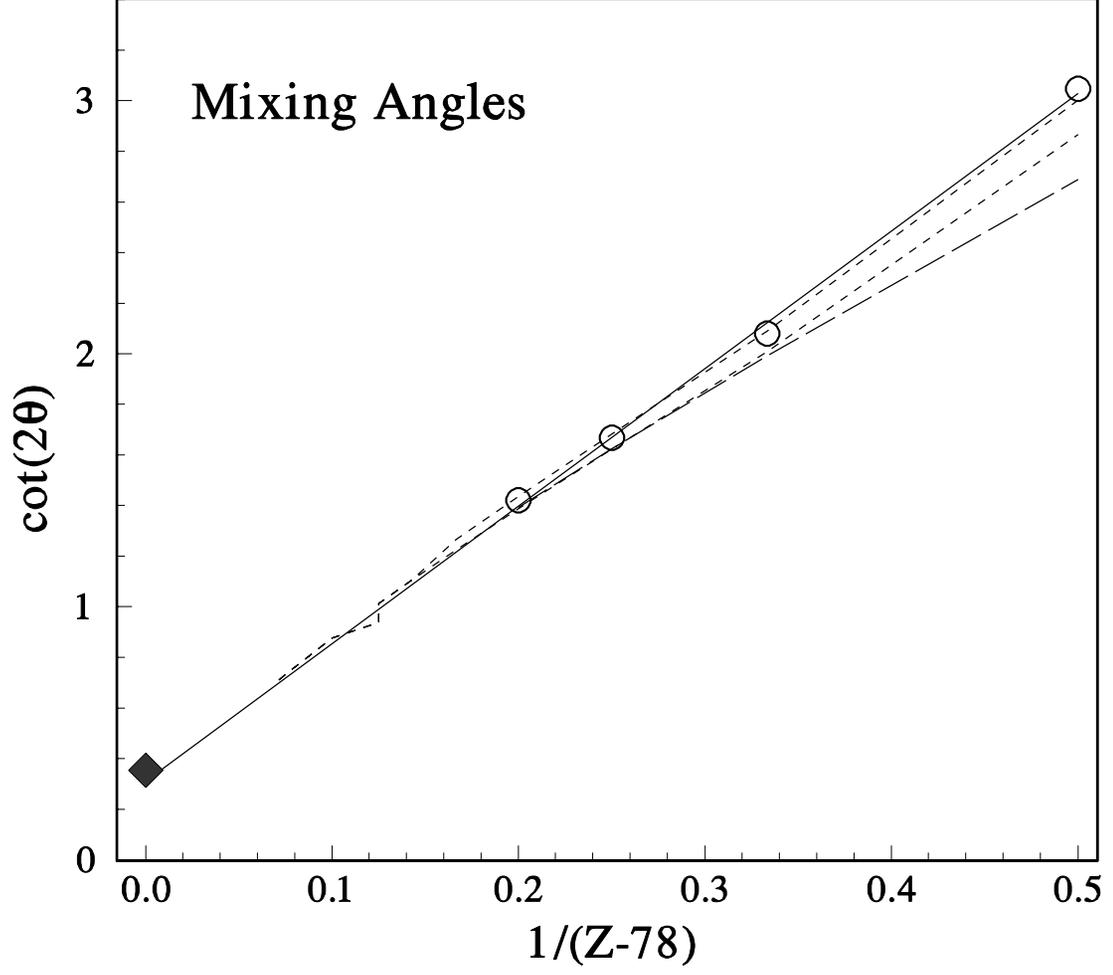
Hg: *80 electrons*
very complex
only four radioactively stable ions

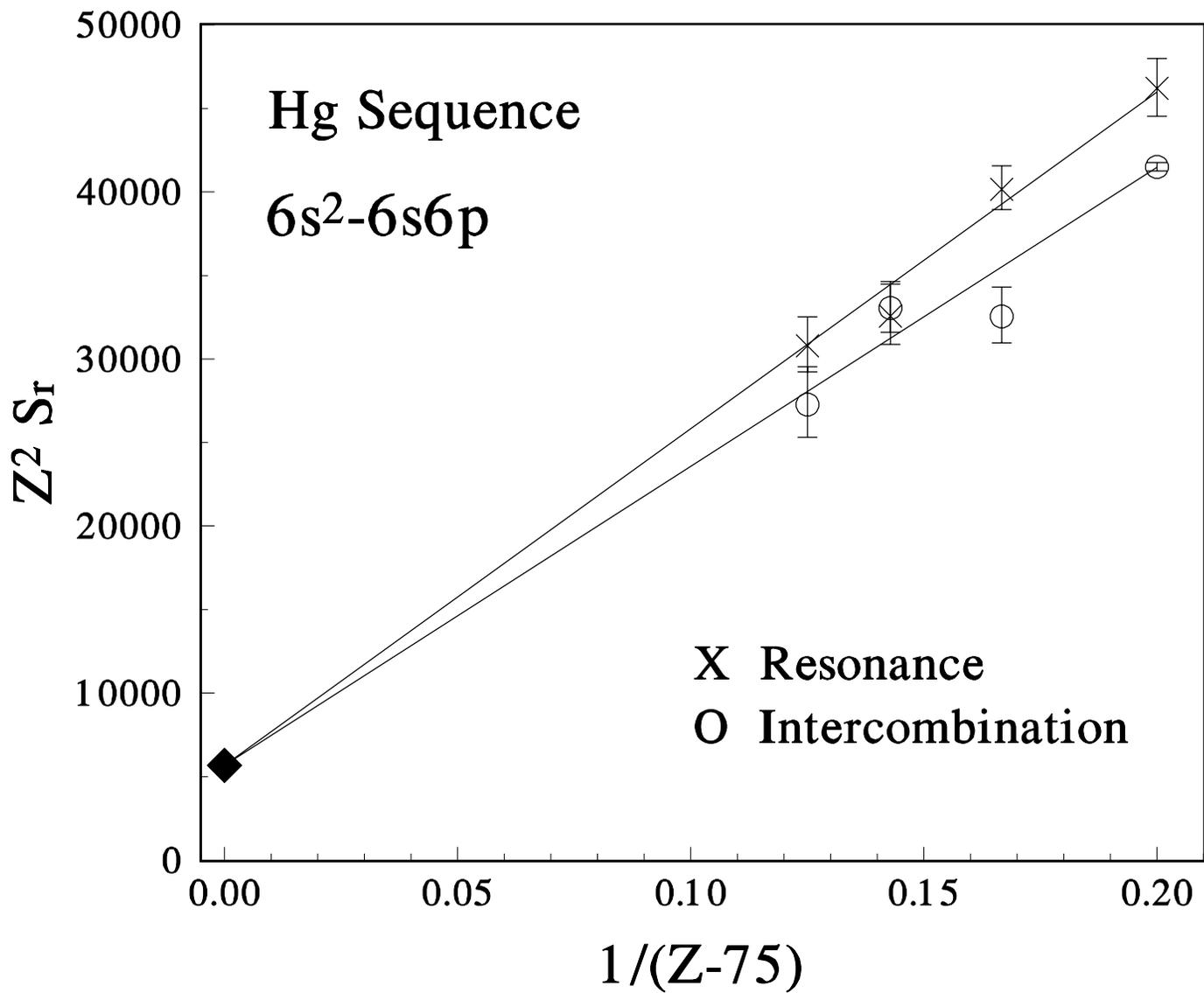
Cd: Plunging levels from unfilled 4f subshell
perturb 5s5p for certain $Z \geq 60$ ions
eventually replace $5s^2$ as ground state

Hg: $6s^2$ and 6s6p remain below levels from 5f
and 5g through $Z = 92$

$6s^2$ -6s6p rates in radioactive elements
difficult to measure, but may dominate
radiative transfer in plasmas that contain
these ions.







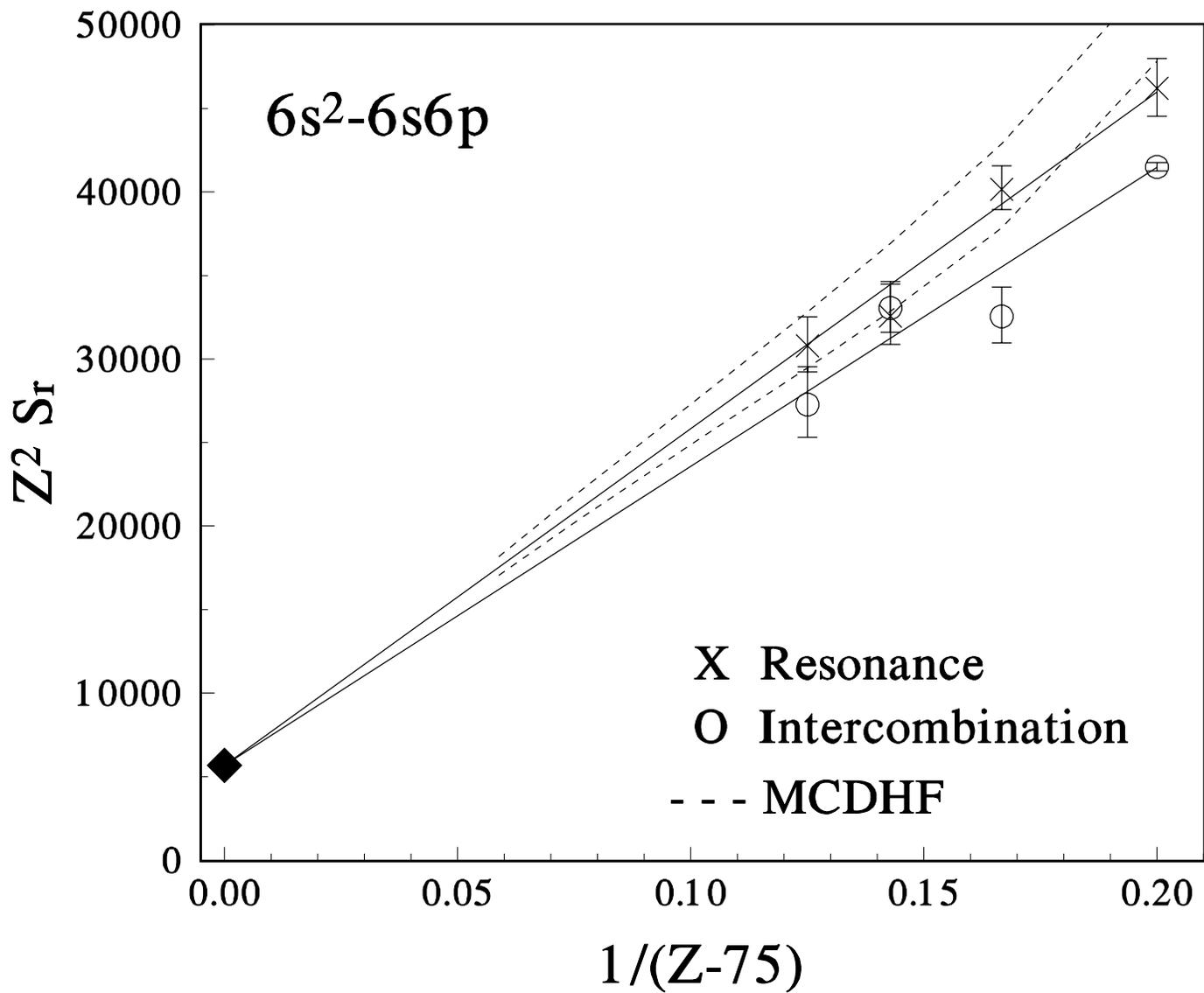


TABLE IV: Lifetime predictions. Parentheses denote quoted or propagated uncertainties.

Z	Ion	Resonance			Intercombination			$\sin(\theta-\xi)^d$		
		$\lambda_{\pi}(A)^a$	$\tau_{\pi}(ns)$	MCDHF ^b	$\lambda_I(A)^a$	$\tau_I(ns)$	MCDHF ^b			
			Expt. ^c	SE ^e	Expt. ^c	SE ^e				
80	Mg I	1049.5	1.54(3) ^a	1.540	1.008	2537.3	118.9(4) ^b	117.9	123.5	0.1771
81	Tl II	1321.7	0.89(4) ^a	0.892	0.543	1908.7	39(3) ^c	36.9	36.0	0.2307
82	Pb III	1048.9	0.300(21) ^d	0.300	0.334	1553.0	15.2(7) ^e	16.1	16.0	0.2725
83	Bi IV	872.6	0.243(13) ^e	0.243	0.228	1317.1	9.1(7) ^e	8.62	8.76	0.3065
84	Po V	748.0		0.176	0.166	1146.3		5.27	5.44	0.3417
85	At VI	653.2		0.123	0.125	1019.4		3.56	3.68	0.3675
86	Rn VII	578.5		0.104	0.0979	919.9		2.57	2.65	0.3899
87	Fr VIII	517.6		0.0829	0.0781	839.2		1.94	2.00	0.4070
88	Ra IX	466.9		0.0671	0.0633	772.0		1.52	1.56	0.4223
89	Ac X	424.0		0.0550	0.0520	715.7		1.23	1.25	0.4355
90	Th XI	387.1		0.0455	0.0432	667.4		1.01	1.03	0.4469
91	Pa XII	355.0		0.0380	0.0362	625.5		0.853	0.865	0.4569
92	U XIII	326.8		0.0319	0.0305	588.7		0.728	0.736	0.4656

^aMCDHF values from Table I for $Z \geq 84$.

^bExtrapolates for $Z \geq 84$ using Eq. 12, 13.

^cRecommended values from Table IV.

^dSemiempirical values from Eqs. 10, 12, 13.

^eCDHF values from this work.

^fSemiempirical values from Eqs. 11, 12, 13.

$$Z^2 S_r(Res) = 5670[1 + 35.55/(Z - 75)]$$

$$Z^2 S_r(Int) = 5670[1 + 31.57/(Z - 75)]$$

$$\cot 2\theta = 0.3096 + 5.437/(Z - 78)$$

$$R_{12}/R_{11} = 0.9931 - 0.0334/(Z - 79)^{1.6}$$

Dirac Formulation

For ns^2 - $nsnp$ there are two radial wavefunctions $R_{2j,2j'}$

$$S(Res) \propto [R_{11} \cos(\theta_{jj} - \theta) - R_{31} \sin(\theta_{jj} - \theta)]^2$$

$$S(Int) \propto [R_{11} \sin(\theta_{jj} - \theta) + R_{31} \cos(\theta_{jj} - \theta)]^2$$

and the jj coupling limit

$$\tan \theta_{jj} = \sqrt{\frac{1}{2}}$$

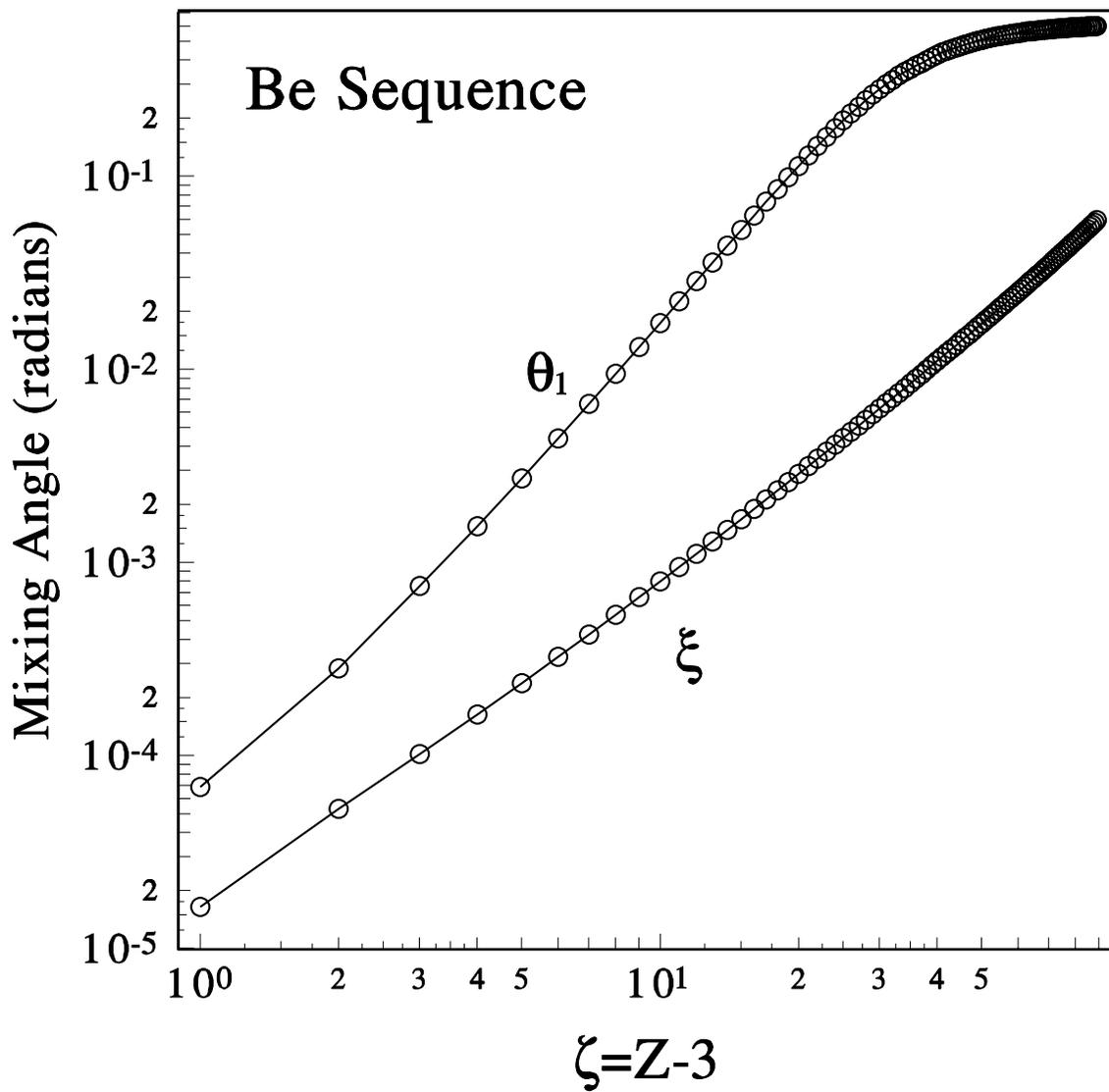
reduces by trigonometric identities to

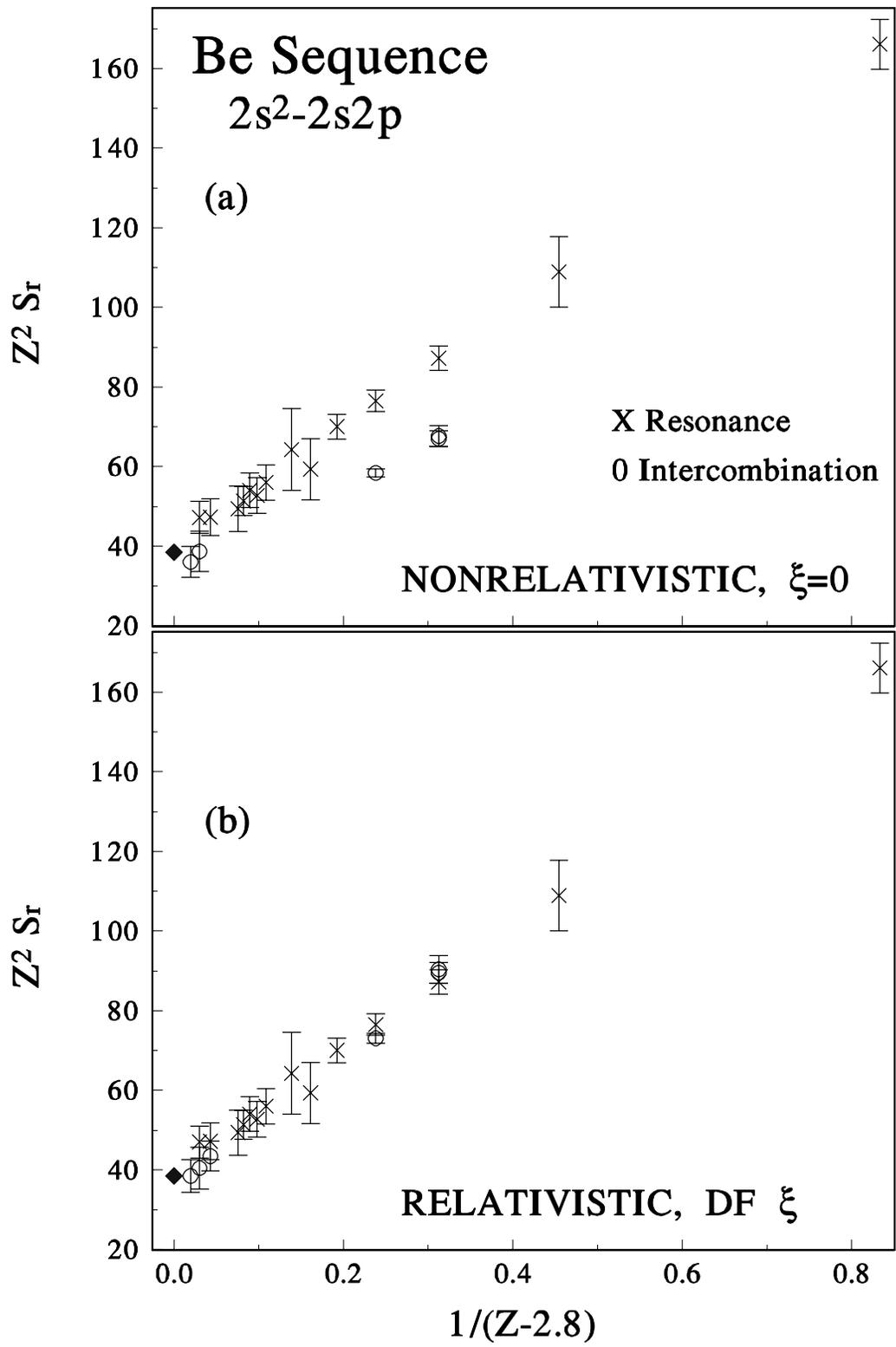
$$S_r(Res) \equiv \frac{S(Res)}{\cos^2(\theta - \xi)}$$

$$S_r(Int) \equiv \frac{S(Int)}{\sin^2(\theta - \xi)}$$

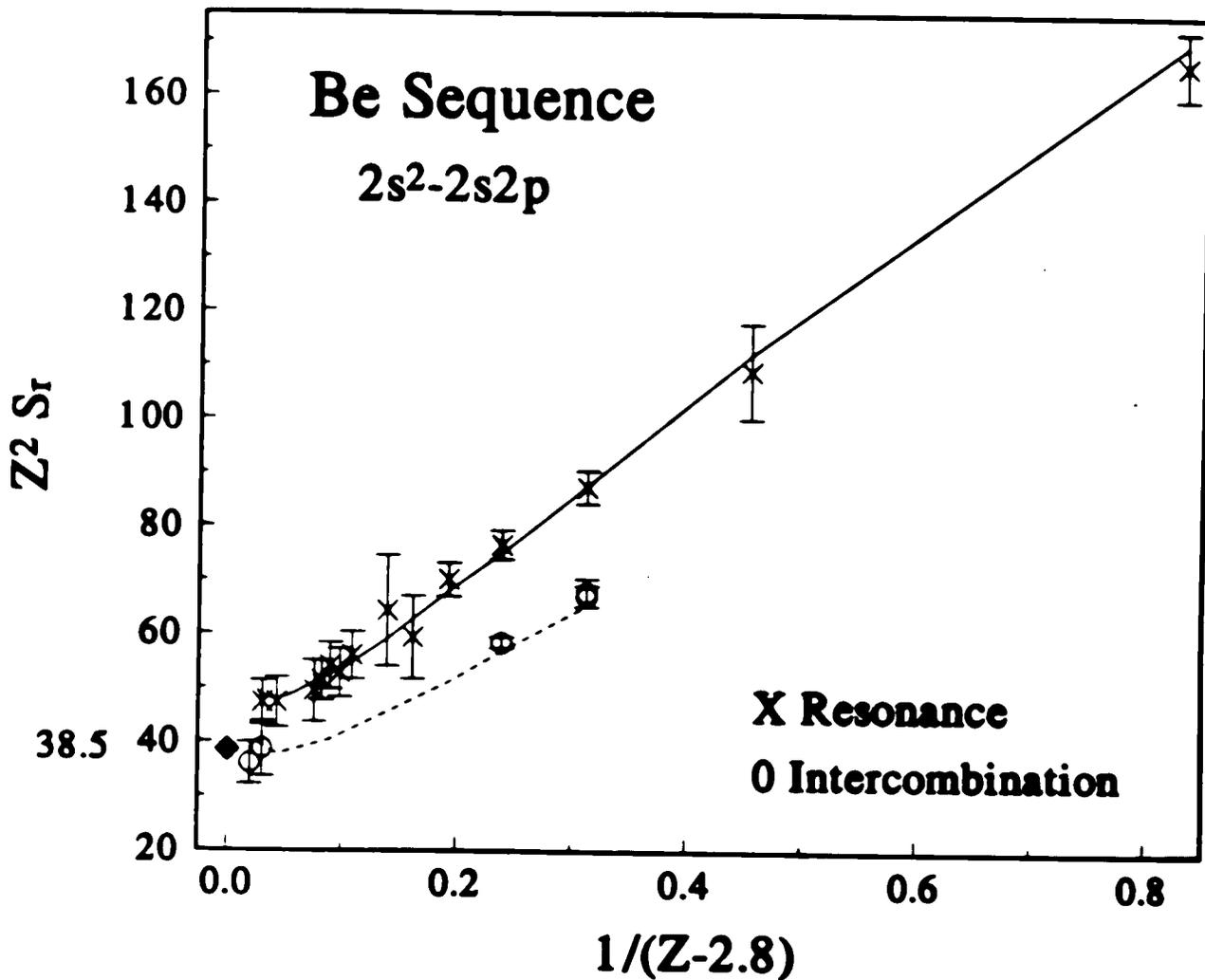
where

$$\tan \xi \equiv \sqrt{2} \frac{R_{31} - R_{11}}{2R_{31} + R_{11}}.$$





IC	$2s2p$	ϑ
CI	$2s^2, 2p^2$	ϕ



$$S_H' = \frac{2}{3} (\sqrt{3} \cos \phi - \sin \phi)^2 S_H \xrightarrow[S_H = S_H', \phi \cong 13^\circ]{} 38.5$$

— Fleming et al, Phys. Scr 53,
446 (1996)

--- Vinnerman & Froese Fischer, PR A
51, 124 (1995)

CLOSED SHELL CORBS WITH HYDROGENIC ORDERING

$$N_e = 2 \sum_{n=1}^{n_{\max}} n^2 = \frac{2}{3} n_n (n_n + \frac{1}{2}) (n_n + 1)$$
$$= 2 (1 + 4 + 9 + 16 + \dots)$$

2 (He-like)
10 (Ne-like)
28 (Ni-like)
60 (Nd-like)

Promethium Isoelectronic Sequence

$$N_e = 61$$

hyper-alkalilike for large Z

Samarium Isoelectronic Sequence

$$N_e = 62$$

hyper-alkaline-earthlike Sequence

He
Ne
Ni
Nd

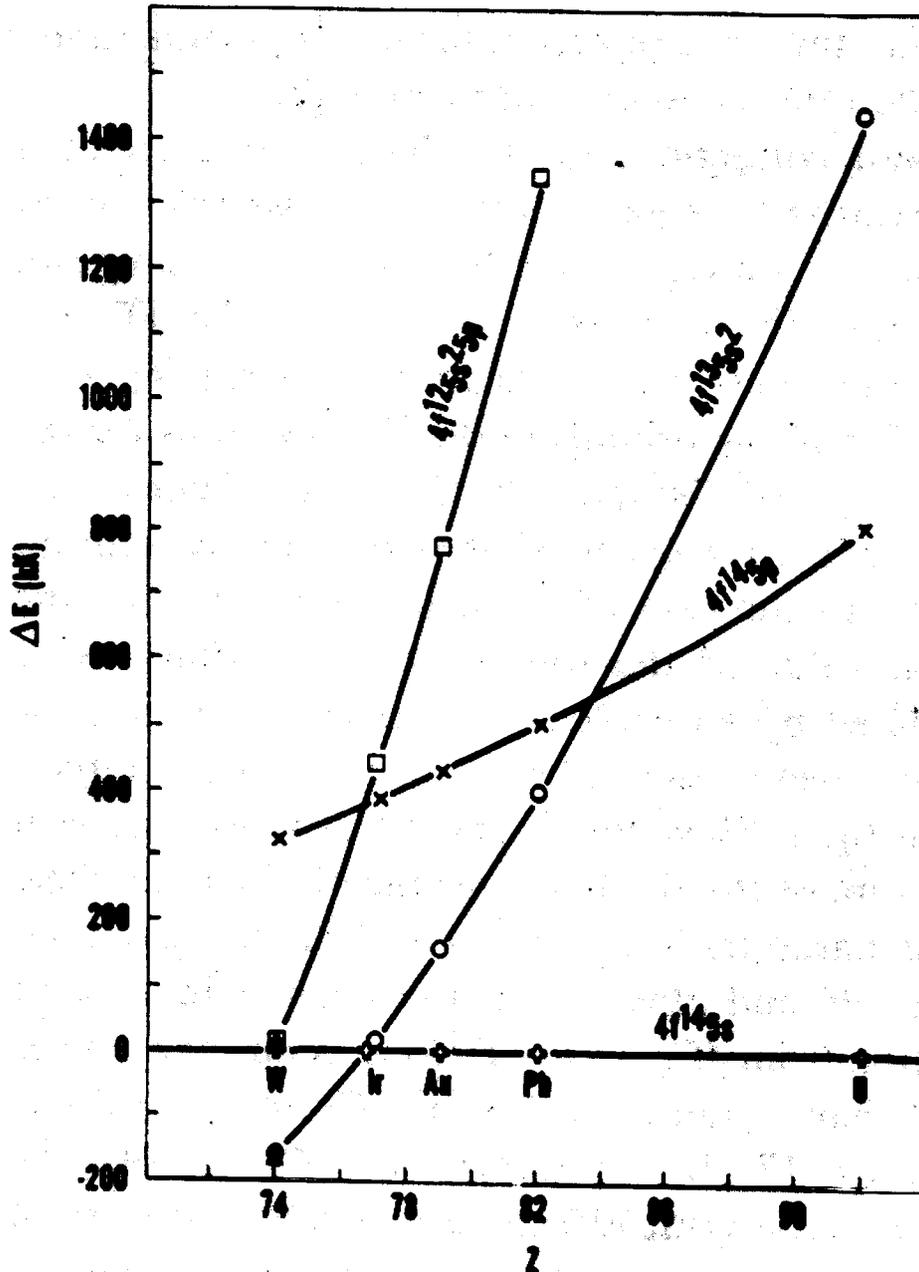
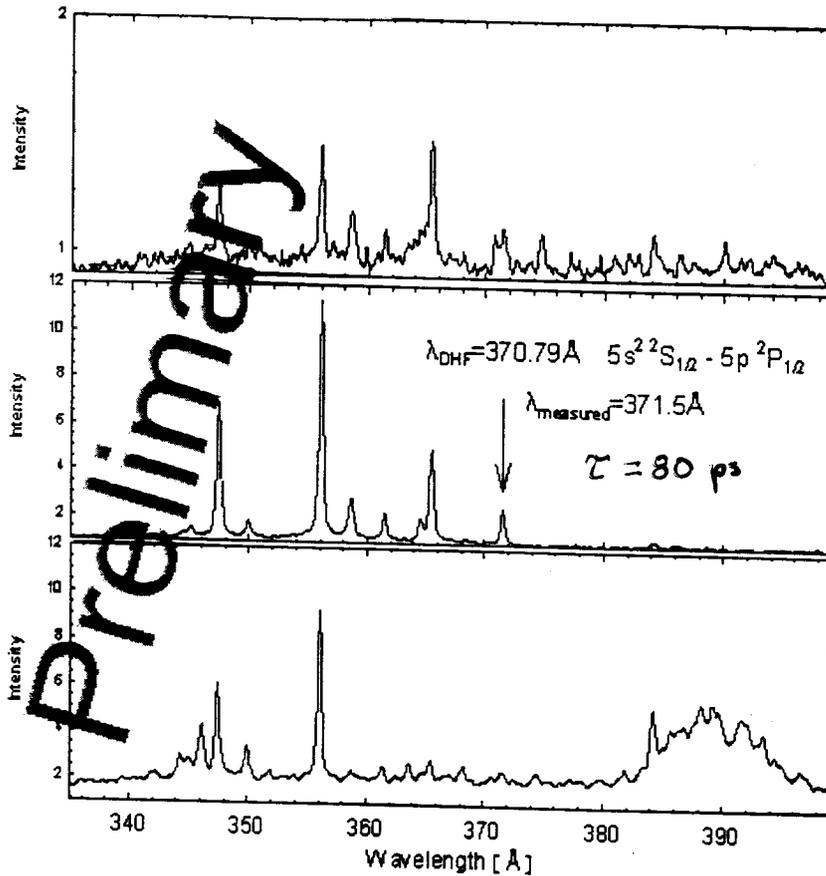


Fig. 1. Upper promethium sequence. The average energies of the indicated configurations ($4f^{14}sp$, $4f^{13}s^2$, $4f^{12}s^2sp$), relative to that of $4f^{14}s$, are plotted against nuclear charge. Included are both the Hartree-Fock energy (HF) and its lowest-order relativistic correction (Pauli). Also plotted for W^{+13} are the relativistic Hartree-Fock (HFR) results of Cowan [5] indicated by triangles. Other symbols indicate results of the present work. Energies are given in units of 10^3 cm^{-1} (kK).

EUV-spectra from tungsten Pm-like W¹³⁺

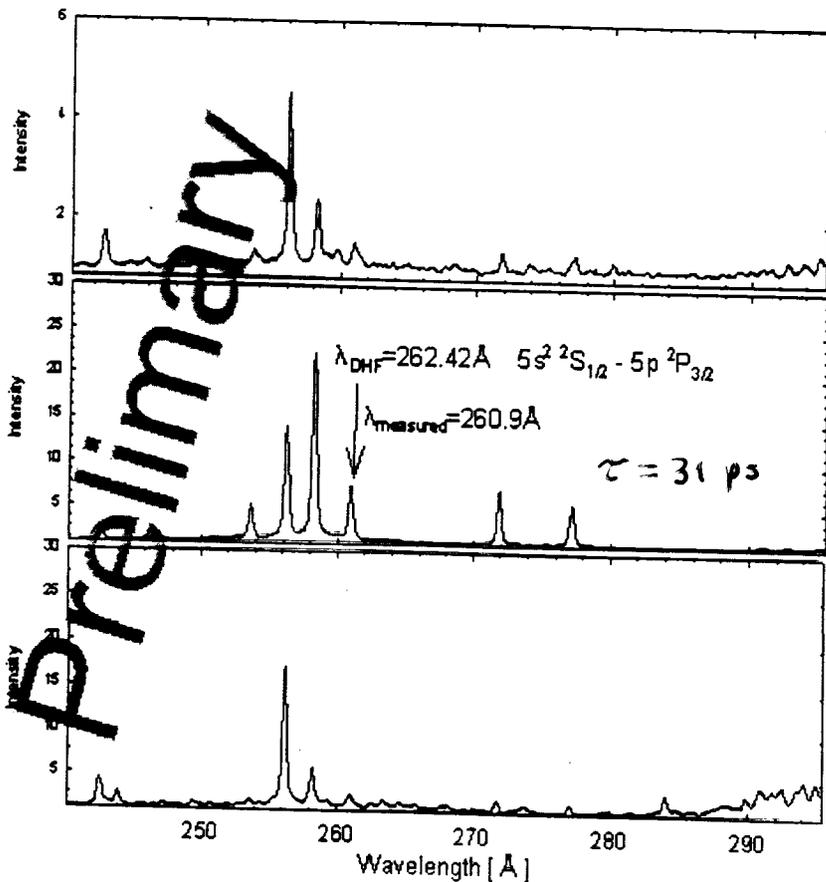


g1018001 $V_{dit} = 0.4000$ kV
 Y=410
 electron beam energy 0.417 keV
 electron beam current 10 mA
 no dump of trap
 data acquisition 60 min
 grating 600 l/mm; 3.5 blaze
 slit 18 μ m

g1018004
 Y = 410
 $V_{dit} = 0.3053$ kV
 electron beam energy 0.311 keV
 electron beam current 10 mA
 60 min

g1018005
 Y = 410
 $V_{dit} = 0.2494$ kV
 electron beam energy 0.244 keV
 electron beam current 10 mA
 60 min

5 m seg
 Res 249 Å 21 ps
 Int 423 Å 752 ps



g1018002
 Y = 348
 $V_{dit} = 0.4000$ kV
 electron beam energy 0.417 keV
 electron beam current 10 mA
 60 min

g1018003
 Y = 348
 $V_{dit} = 0.3053$ kV
 electron beam energy 0.311 keV
 electron beam current 10 mA
 ca. 40 min

g1018007
 Y = 348
 $V_{dit} = 0.2475$ kV
 electron beam energy 0.242 keV
 electron beam current 10 mA
 30 min

Projection within 385+250 pixels
 calib. Sep2000

SAMARIUM SEQUENCE									
Z Ion	1	1			1	3			
	(S	-	P)		(S	-	P)		
	0	1			0	1			
	λ (Å)	τ (ps)			λ (Å)	τ (ps)			
	MCDF	MCDF	Mix		MCDF	Obs	MCDF	Mix	HF/CI Obs
74	249.3	20.7			423.3		752		
75 Re	237.5	17.4	11.9		398.4		608	393	
76 Os	221.3	12.8	10.3		376.2	[366.3]	502	333	281
77 Ir	207.0	11.1	9.0		356.3	[347.0]	421	285	240 [260]
78 Pt	193.9	9.7	8.1		338.3	[330.0]	357	248	207 [200]
79 Au	182.0	8.6	7.3		321.8	[313.6]	307	219	180 [170]
80 Hg	171.0	7.5	6.5		306.8		266	194	
81 Tl	160.9	6.7	5.8		293.1		232	172	
82 Pb	151.5	5.9	5.1		280.4		205	153	
83 Bi	142.8	5.2	4.5		268.7		182	136	
84 Po	134.7	4.6	4.0		257.8		162	120	
85 At	127.2	4.1	3.5		247.7		146	107	
86 Rn	120.2	3.7	3.1		238.2		132	95	
87 Fr	113.6	3.3	2.8		229.4		119	86	
88 Ra	107.5	2.9	2.5		221.2		109	79	
89 Lu	101.8	2.6	2.2		213.4		100	74	
90 Th	96.4	2.3	2.1		206.1		92	71	
91 Pa	91.3	2.1	2.0		199.3		84	70	
92 U	86.6	1.8	1.9		192.8		78	71	

MCDF: Curtis, JOSA B3, 1102 (1986); Grant-2 program.
 Mix : Singlet-Triplet mixing with Pm line strengths.
 Pm from Curtis & Ellis, PRL 45, 2099 (1980); MCDF+BP.
 HF/CI: Kaufman (person communication); Cowan code.
 Obs: Kaufman, Heckmann, Träbert, Möller, Ludwig & Blank
 (personal communication). Phys. Scr. 42, 705 (1990).

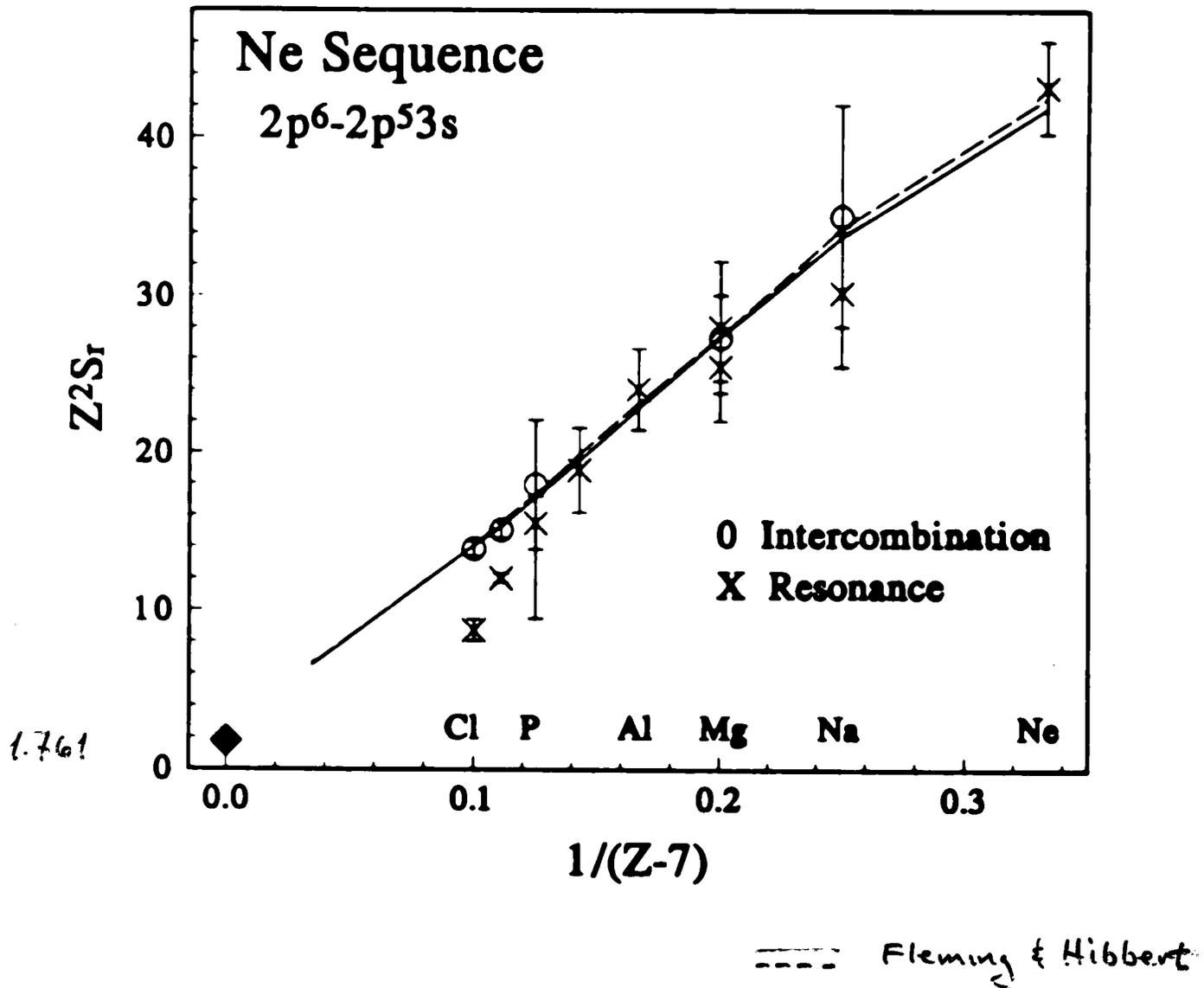


TABLE III. Ne sequence $2p^5 3s^1, ^3P$ levels (lifetimes in ps).

Z	Ion	Intercombination transitions		Resonance transitions		sin θ
		τ_{expt}	τ_{pred}	τ_{expt}	τ_{pred}	
10	Ne	$29600 \pm 1000, 10^{b,j}$ $31700 \pm 1600, 10^{b,j}$ $29800 \pm 2000, 10^{c,j}$		1470 ± 100^a $1870 \pm 180, 10^{b,j}$ $1300 \pm 100, 10^{c,j}$	1547	0.2656
11	Na	6000 ± 1200^a $10600 \pm 500, 10^{d,j}$	5623	320^{+50a}_{-40} 580 ± 60^d	346	0.2131
12	Mg	1900 ± 190^e	1820	100 ± 15^e 110^{+15a}_{-10}	131	0.2299
13	Al		637	46^{+10a}_{-5}	63	0.2677
14	Si		248	28^{+6a}_{-4}	36	0.3171
15	P	130 ± 30^a	137	18^{+7a}_{-2}	21	0.3731
16	S	52 ± 2^f $49 \pm 13^{h,j}$	52	14.5 ± 1.0^g $12 \pm 3^{h,j}$	15	0.4309
17	Cl	27 ± 1^f $34 \pm 12^{h,j}$ $30 \pm 5^{i,j}$	27	13 ± 1^f $10 \pm 2^{h,j}$ $8 \pm 2^{i,j}$	11	0.4872
18	Ar	$19 \pm 4^{i,j}$	16	$6.5 \pm 2.0^{i,j}$		
34	Se		0.28		0.60	0.7847
35	Br		0.24		0.53	0.7878
36	Kr		0.21		0.47	0.7904
37	Rb		0.19		0.42	0.7927
38	Sr		0.16		0.36	0.7947
39	Y		0.14		0.34	0.7966

^aThis work.

^bLawrence and Liszt, Ref. [56].

^cKernahan *et al.*, Ref. [57].

^dSchlagheck, Refs. [58,59].

^eBuchet *et al.*, Ref. [60].

^fWesterlind *et al.*, Ref. [61].

^gKirm *et al.*, Ref. [65].

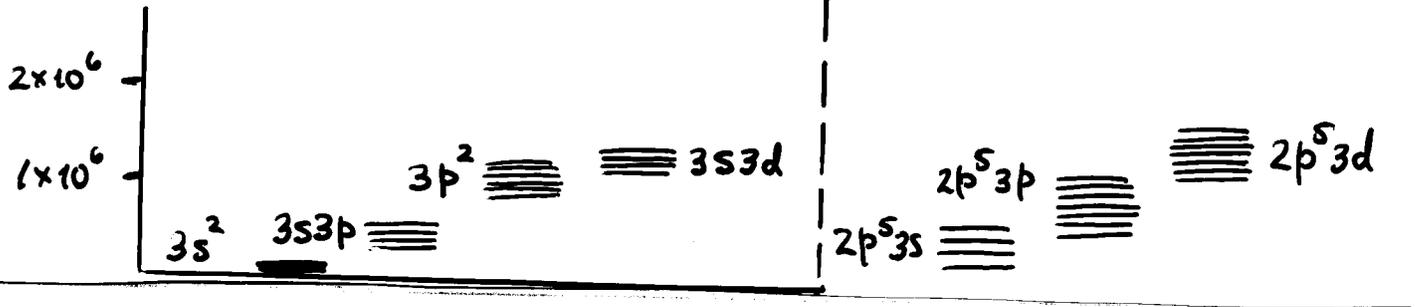
^hGardner *et al.*, Ref. [62].

ⁱBerry *et al.*, Ref. [63].

^jExcluded from plot and fit.

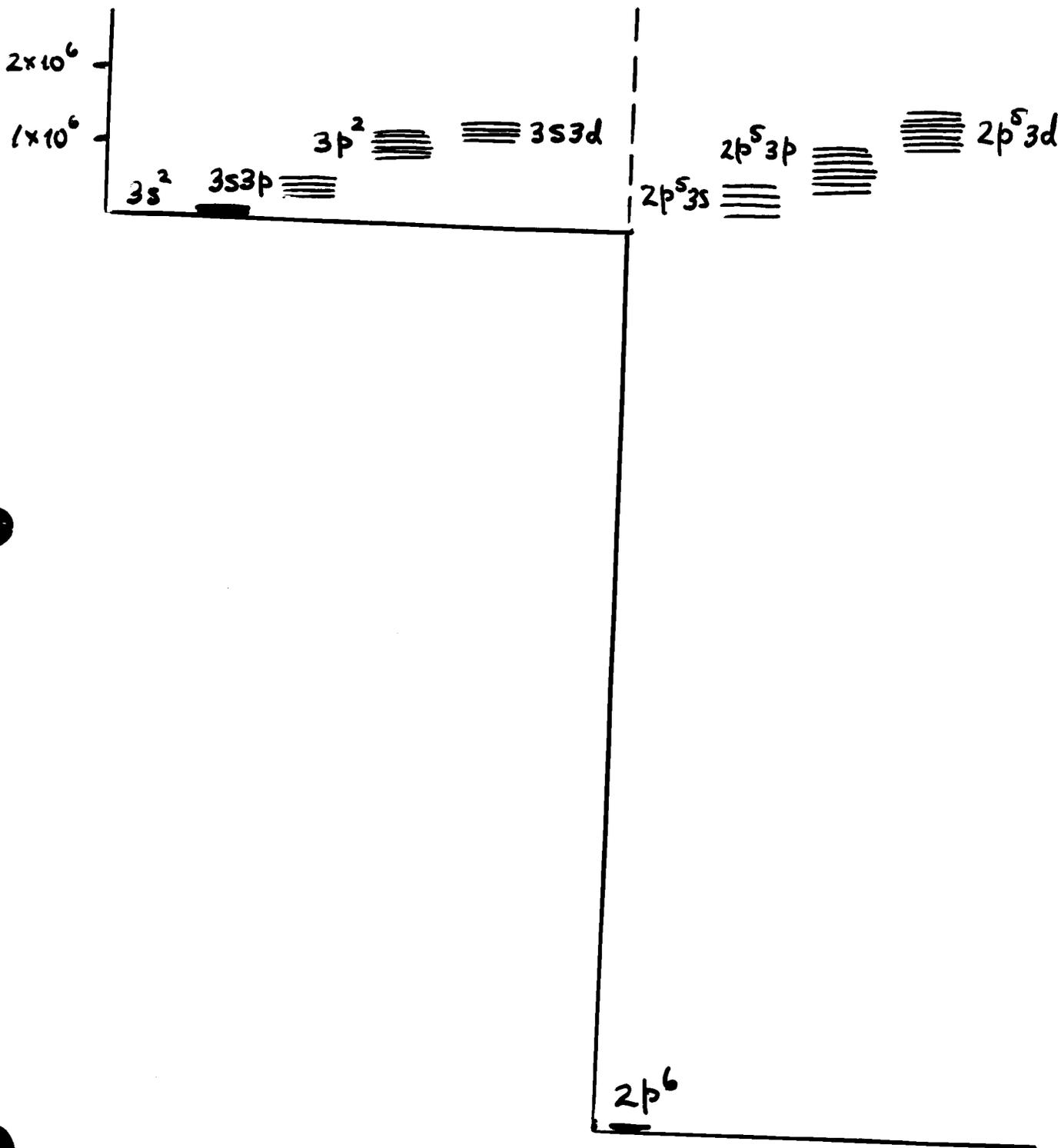
Mg-like Br XXIV

Ne-like Br XXVI

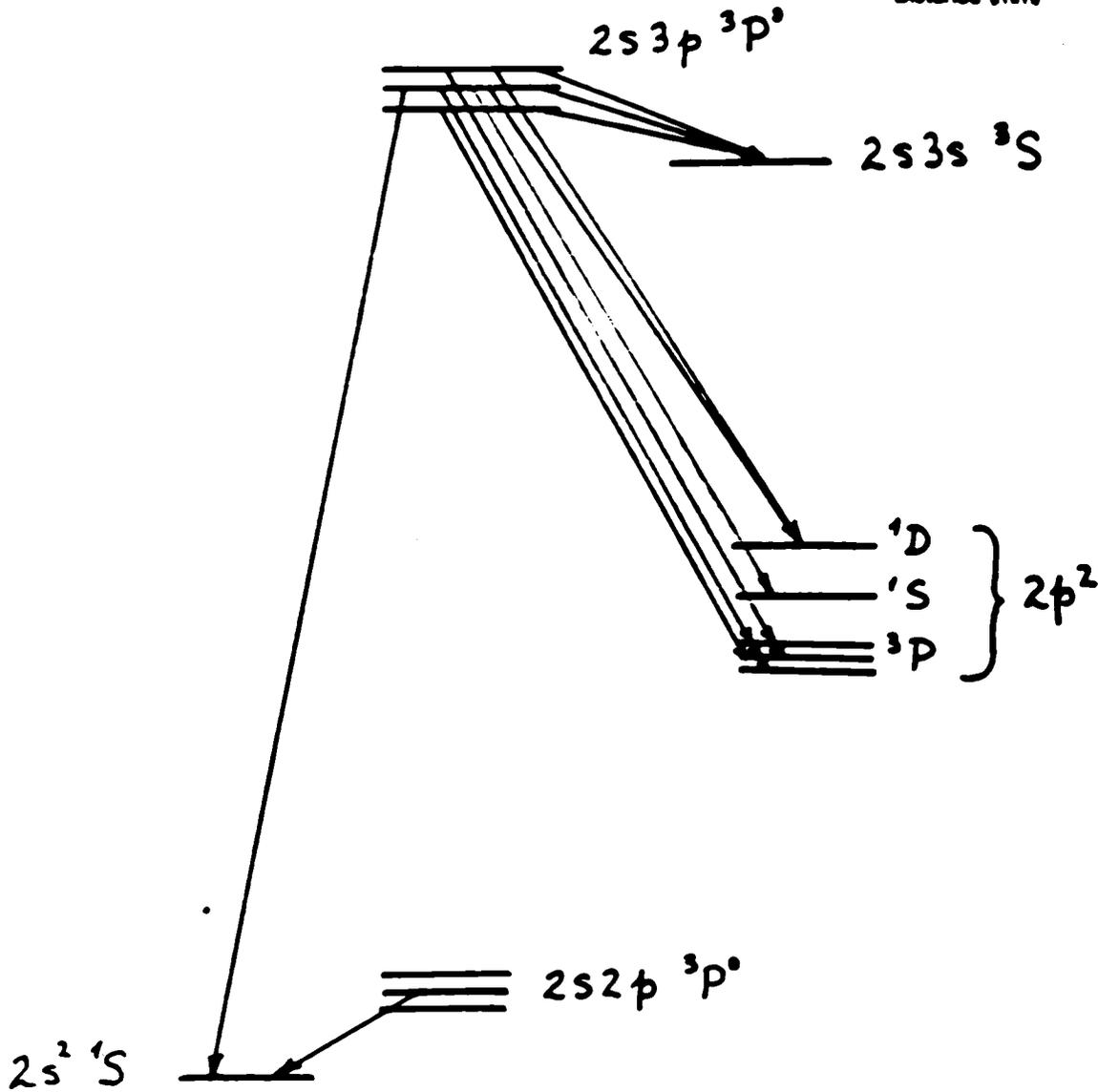
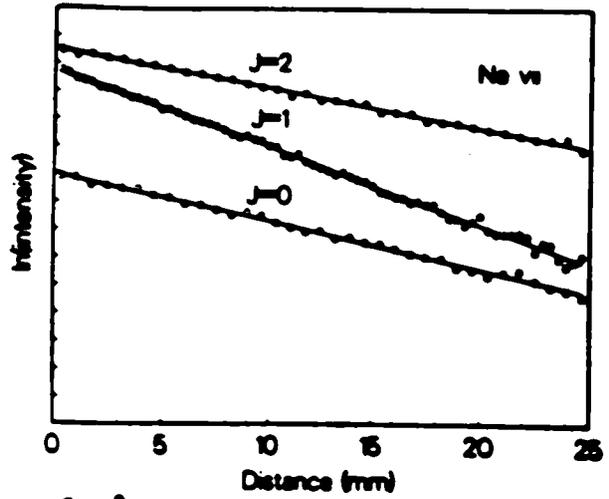


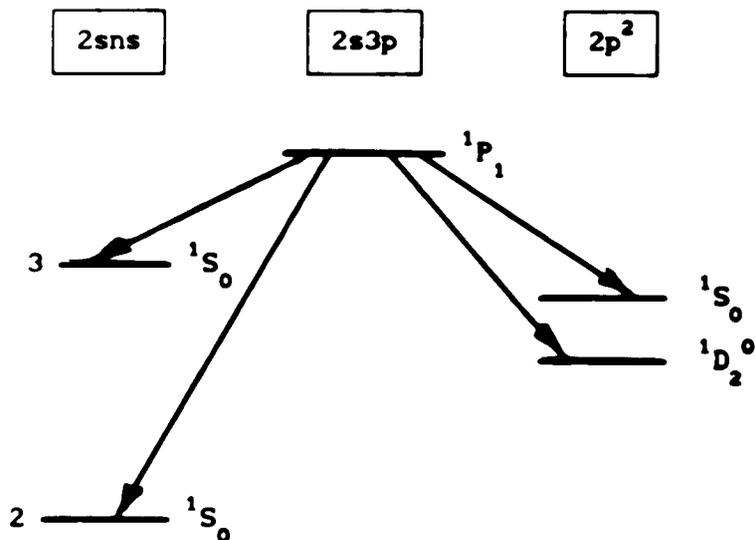
Mg-like Br XXIV

Ne-like Br XXVI



BRANCHING FRACTIONS FROM DIFFERENTIAL LIFETIME MEASUREMENTS



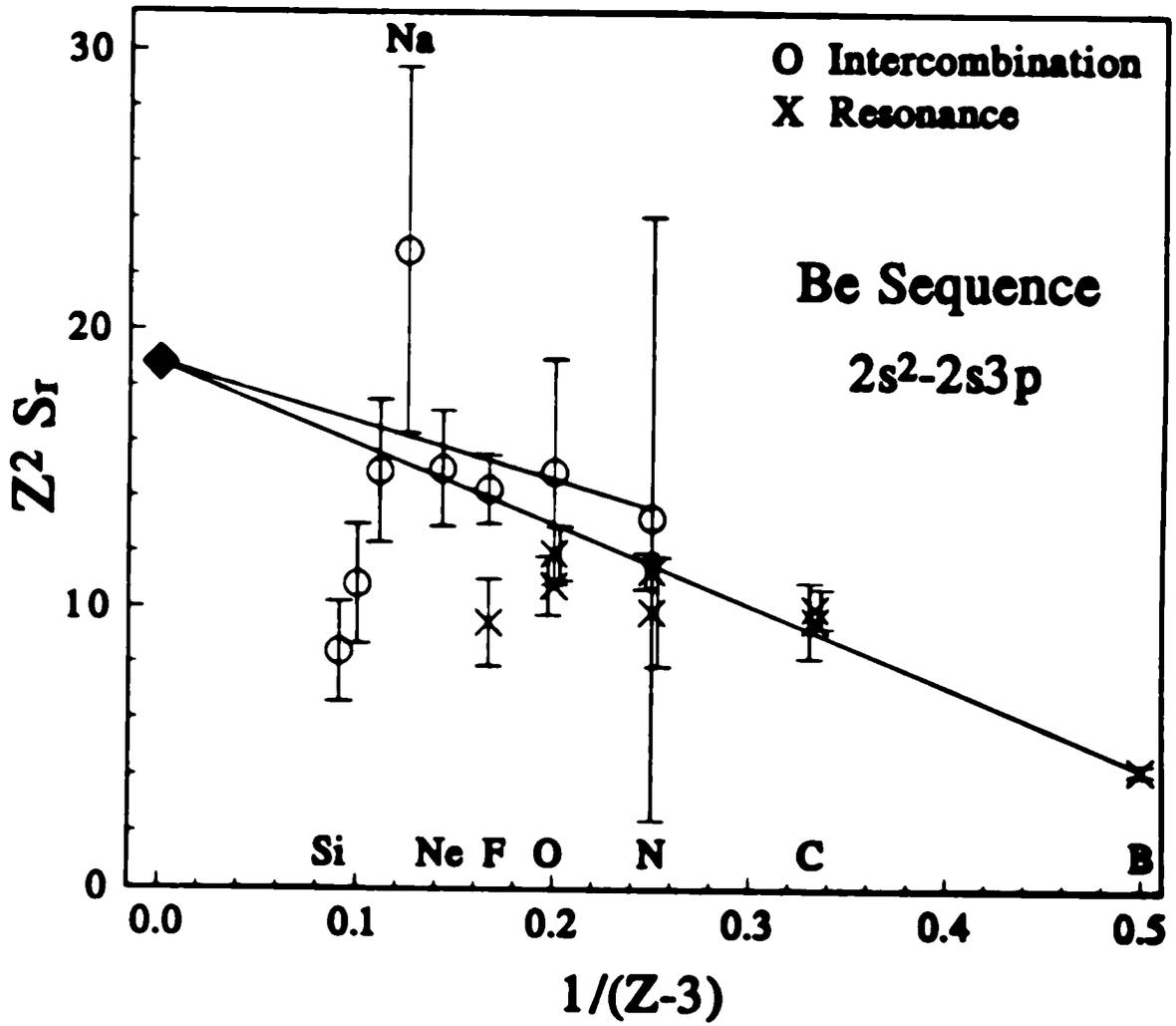


Branching Fractions (X)

$2s3p$ $^1P_1 \rightarrow$	$2s^2$ 1S_0	$2p^2$ 1D_2	$2p^2$ 1S_0	$2s3s$ 1S_0
Be I	55.6	26.5	0.1	17.9
B II	68.6	14.4	13.9	3.0
C III	87.3	11.6	0.8	0.2
N IV	89.9	9.5	0.4	0.1
O V	91.4	8.2	0.3	0.05
Ne VII	93.5	6.3	0.2	0.02
Fe XXIII	95.9	3.1	0.1	0.01

Bhatia & Mason, A&A 103, 324 (1981) [Fe]

Laughlin et al, J.Phys.B 11, 2243 (1978) [Others]



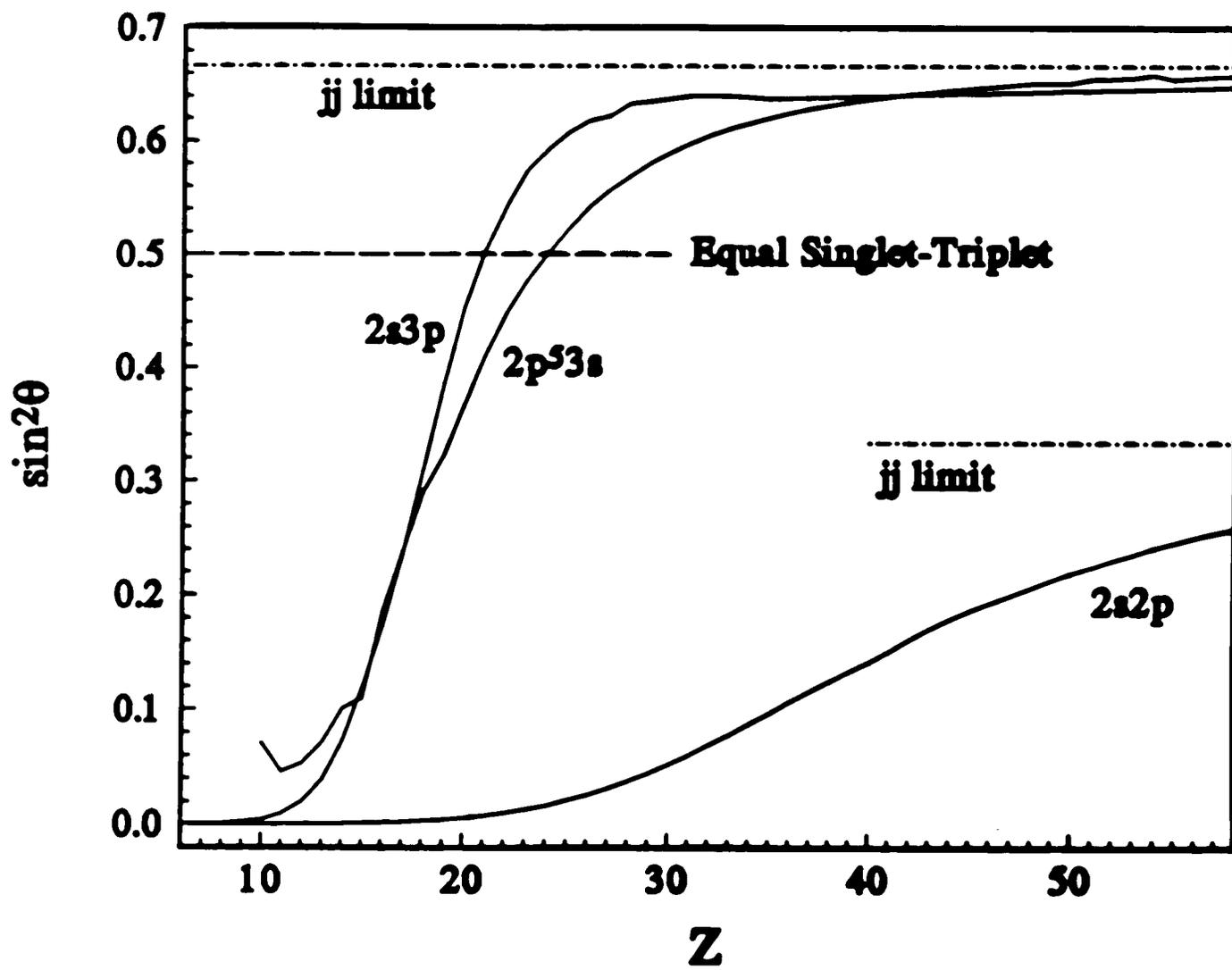


TABLE II. Be sequence $2s3p\ ^1,^3P$ levels (lifetimes in ns). The experimental uncertainty in the last figure is given in parentheses. Branching fractions without labels are interpolated.

Z	Ion	Intercombination transitions				Resonance transitions				sin θ
		$\tau_{2,0}$	τ_1	A_{expt}	A_{pred}	τ	BF	A_{expt}	A_{pred}	
5	B				0.0001	1.98(9) ^a	0.696 ^b	0.346(16)	0.38	0.0199
6	C				0.0004	0.28(4) ^c	0.873 ^b	3.12(45)	2.96	0.0113
7	N	8.98(10)	8.72(12)	0.0033(20) ^a	0.0028	0.268(20) ^d	0.809 ^b	3.26(24)		
8	O	5.31(5)	4.74(7)	0.0226(4) ^a	0.019	0.067(5) ^c		10.3(6)	10.4	0.0164
						0.10(2) ^e		9(2)		
9	F	2.94(3)	2.24(4)	0.102(9) ^c	0.100	0.041(4) ^c	0.914 ^b	22(2)	26.5	0.0206
10	Ne	1.75(7)	1.00(6)	0.43(6) ^b	0.418	0.637(3) ^c		25(2)		
11	Na	1.10(16)	0.29(6)	2.54(73) ⁱ	1.68	0.824(4) ^c	0.927	30(6)	56.2	0.0419
12	Mg	0.745(60)	0.145(20)	5.55(96) ⁱ	5.78		0.934 ^b		105	0.0327
13	Al	0.515(50)	0.071(12)	12.1(24) ⁱ	17.6		0.930		180	0.0060
14	Si	0.365(25)	0.036(7)	25.0(54) ⁱ	47.7		0.943		208	0.1400
15	P				111		0.946		430	0.1900
16	S				223		0.949		600	0.3008
17	Cl				410		0.951		822	0.3441
18	Ar				711		0.952		1085	0.4155
19	K				1153		0.953		1327	0.4851
20	Ca				1732		0.955		1594	0.5558
21	Sc				2426		0.955		1828	0.6200
22	Ti				3226		0.956		2079	0.6732
23	V				4207		0.957		2375	0.7098
24	Cr				5130		0.957		2740	0.7390
25	Mn				6303		0.958		3151	0.7574
26	Fe				7850		0.958		3636	0.7690
							0.959 ^f		4086	0.7792
									5130	0.7980

^aKernahan *et al.*, Ref. [57].

^bLaughlin *et al.*, Ref. [40].

^cThis work.

^dBuchet-Poulizac and Buchet, Ref. [37].

^eEngström *et al.*, Ref. [27].

^fDumont *et al.*, Ref. [38].

^gKnystantes and Drouin, Ref. [39].

^hHardis *et al.*, Ref. [36].

ⁱGrassew *et al.*, Ref. [28].

^jBhatia and Mason, Ref. [41].

For the s^2 - sp transitions these are

$$\langle {}^1S_0 | r | {}^3P_1 \rangle = [\sin \theta_1] \mathcal{M}_{sp}$$

$$\langle {}^1S_0 | r | {}^1P_1 \rangle = [\cos \theta_1] \mathcal{M}_{sp}$$

and for sp - sd

$$\langle {}^3P_0 | r | {}^3D_1 \rangle = \left[\frac{1}{\sqrt{3}} \right] \mathcal{M}_{pd},$$

$$\langle {}^3P_1 | r | {}^3D_1 \rangle = \left[\frac{1}{2} \cos \theta_1 \right] \mathcal{M}_{pd},$$

$$\langle {}^3P_2 | r | {}^3D_1 \rangle = \left[\frac{1}{\sqrt{60}} \right] \mathcal{M}_{pd},$$

$$\langle {}^1P_1 | r | {}^3D_1 \rangle = \left[\frac{1}{2} \sin \theta_1 \right] \mathcal{M}_{pd},$$

$$\langle {}^3P_1 | r | {}^3D'_2 \rangle = \left[\frac{\sqrt{3}}{2} \cos \theta_1 \cos \theta_2 + \sin \theta_1 \sin \theta_2 \right] \mathcal{M}_{pd},$$

$$\langle {}^3P_2 | r | {}^3D'_2 \rangle = \left[\frac{1}{2} \cos \theta_2 \right] \mathcal{M}_{pd},$$

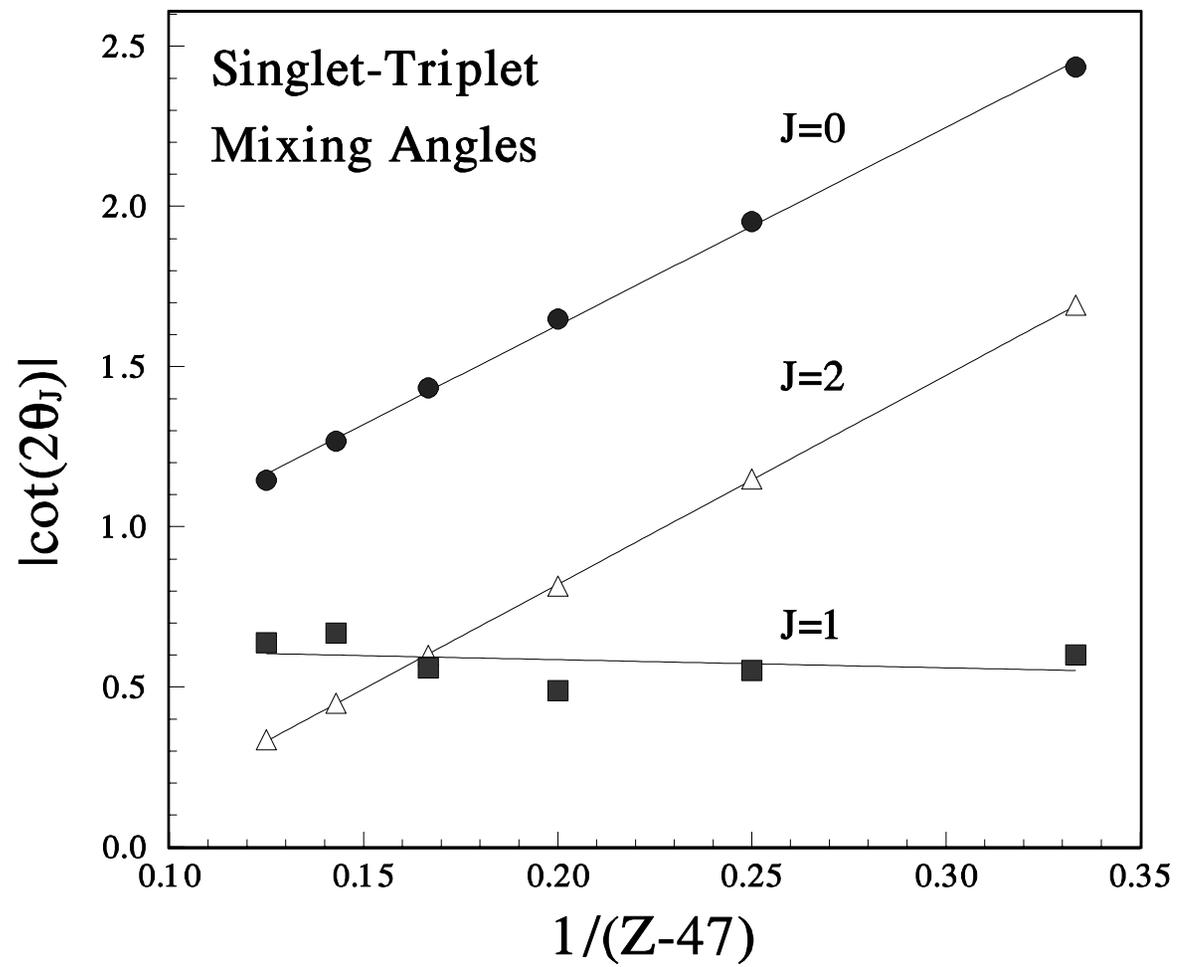
$$\langle {}^1P_1 | r | {}^3D'_2 \rangle = \left[\frac{\sqrt{3}}{2} \sin \theta_1 \cos \theta_2 - \cos \theta_1 \sin \theta_2 \right] \mathcal{M}_{pd},$$

$$\langle {}^3P_2 | r | {}^3D_3 \rangle = \left[\sqrt{\frac{7}{5}} \right] \mathcal{M}_{pd},$$

$$\langle {}^3P_1 | r | {}^1D'_2 \rangle = \left[\frac{\sqrt{3}}{2} \cos \theta_1 \sin \theta_2 - \sin \theta_1 \cos \theta_2 \right] \mathcal{M}_{pd},$$

$$\langle {}^3P_2 | r | {}^1D'_2 \rangle = \left[\frac{1}{2} \sin \theta_2 \right] \mathcal{M}_{pd},$$

$$\langle {}^1P_1 | r | {}^1D'_2 \rangle = \left[\frac{\sqrt{3}}{2} \sin \theta_1 \sin \theta_2 + \cos \theta_1 \cos \theta_2 \right] \mathcal{M}_{pd}.$$



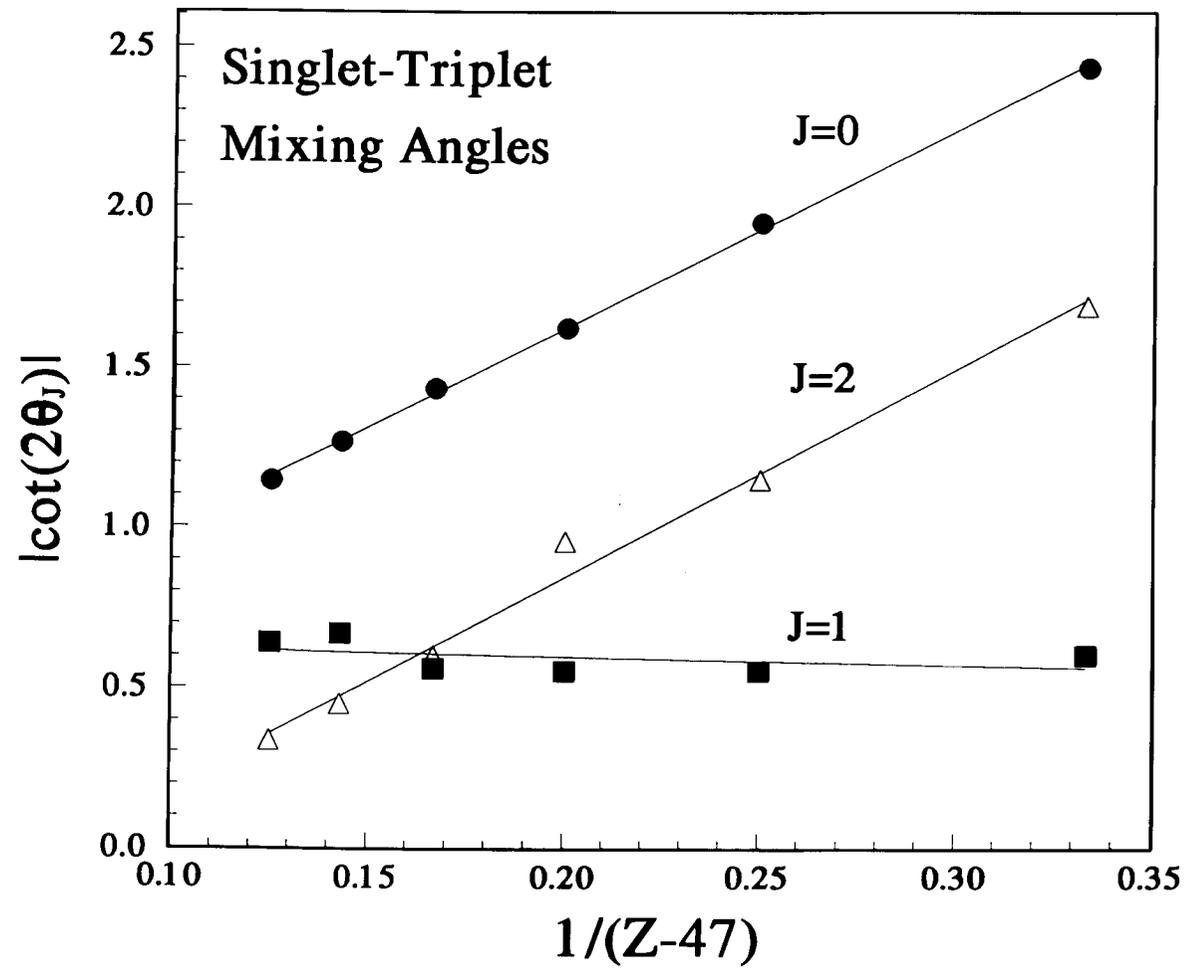


Table 4: Data base of experimental lifetime measurements (Exp) and semiempirical predictions (SE) based on this parametrization. Quoted measurement uncertainties are given in parentheses. The 3P_1 predictions are based on the 1P_1 measurements, and the 3D_1 and 3D_3 predictions are based on the 3D_2 measurements. A few recent calculations for the lifetimes of the $nsnp$ levels (Theo) are included for comparison.

Level	Ga II			In II			Tl II		
	Exp	Theo	SE ^a	Exp	Theo	SE ^a	Exp	Theo	SE ^a
3P_1	-	2445 ^b	2380	440(40) ^c	598 ^d	449	39(3) ^e	36.3 ^f	34.2
1P_1	0.41(3) ^g , 0.49(4) ^h , 0.65(8) ⁱ , 0.48(12) ^j	-	-	0.79(5) ^k , 0.90(8) ^l	-	-	0.59(4) ^e	0.574 ^f	-
3D_1	-	-	0.66	0.86(3) ^k	-	0.86	-	-	-
3D_2	0.67(6) ^j	-	-	0.91(3) ^k	-	-	-	-	-
3D_3	-	-	0.69	0.94(3) ^k	-	1.00	-	-	-
1D_2	0.67(4) ^j , 0.73(7) ^m	-	-	0.77(3) ^k	-	-	5(1) ⁿ , 7(1) ^o	-	-

^a This work.

^b Fleming and Hibbert [19].

^c Peik *et al.* [22].

^d Chou *et al.* [20].

^e Henderson and Curtis [23].

^f Brage *et al.* [21].

^g Engström [24].

^h Andersen *et al.* [25].

ⁱ Sørensen [26].

^j Ansbacher *et al.* [27].

^k Ansbacher *et al.* [28].

^l Andersen *et al.* [29].

^m Denne *et al.* [30].

ⁿ Andersen and Sørensen [31].

^o Shimon and Erdevdi [32].

Table 2: Wavelengths (in air for $\lambda \geq 2000$ Å), multiplet fractions (in %), branching fractions (in %) and transition probability rates (in ns^{-1}) for the two supermultiplets. The transition probability rate predictions are based on the branching fractions obtained by this formalism and the measured and predicted lifetimes given in Table 3.

Transition	Ga II				In II				Tl II			
	$\lambda(\text{Å})$	R_{ik}	BF	A_{ik}	$\lambda(\text{Å})$	R_{ik}	BF	A_{ik}	$\lambda(\text{Å})$	R_{ik}	BF	A_{ik}
$^1\text{S}_0$ - $^3\text{P}_1$	2090.77	0.019	100	0.00042	2306.15	0.18	100	0.0023	1908.65	1.70	100	0.026
$^1\text{S}_0$ - $^1\text{P}_1$	1414.40	99.98	100	2.19	1586.45	99.82	100	1.22	1321.70	98.30	100	1.69
$^3\text{P}_0$ - $^3\text{D}_1$	1504.93	15.94	56.13	0.854	1672.00	16.57	57.17	0.665	1499.34	4.14	60.31	-
$^3\text{P}_1$ -	1515.11	11.71	41.23	0.627	1702.57	11.70	40.38	0.470	1568.53	8.81	37.55	-
$^3\text{P}_2$ -	1536.90	0.75	2.63	0.040	1777.57	0.69	2.38	0.028	1837.49	0.38	1.64	-
$^1\text{P}_1$ -	2318.68	0.002	0.01	0.0001	2560.06	0.02	0.07	0.001	2469.18	0.12	0.50	-
SUM		28.40	100	1.522		28.98	100	1.163		23.45	100	-
$^3\text{P}_1$ - $^3\text{D}_2$	1514.51	21.11	75.79	1.131	1700.08	21.20	77.22	0.849	1561.60	25.22	79.30	-
$^3\text{P}_2$ -	1536.28	6.74	24.21	0.361	1774.86	6.23	22.70	0.249	1827.99	5.80	18.24	-
$^1\text{P}_1$ -	2317.27	0.002	0.01	$<10^{-4}$	2554.44	0.023	0.08	0.001	2452.04	0.78	2.46	-
SUM		27.85	100	1.493		27.46	100	1.099		31.80	100	-
$^3\text{P}_2$ - $^3\text{D}_3$	1535.31	27.03	100	1.448	1770.66	25.11	100	1.063	1814.85	33.49	100	-
$^3\text{P}_1$ - $^1\text{D}_2$	1275.94	0.16	0.02	0.0003	1417.81	0.20	1.08	0.014	1593.19	3.21	28.52	0.048
$^3\text{P}_2$ -	1291.36	$<10^{-3}$	$<10^{-3}$	10^{-5}	1469.44	0.002	0.01	0.0001	1871.43	0.05	0.44	0.001
$^1\text{P}_1$ -	1802.25	16.70	99.98	1.460	1966.71	18.25	98.91	1.285	2530.88	7.99	71.04	0.118
SUM		16.72	100	1.460		18.45	100	1.299		11.25	100	0.167

TABLE III. Comparison of predicted transition probabilities and branching fractions for the $5s5p\ ^3P - 5s5d\ ^3D$ manifold in In II.

Transition	$A_{ik}(ns^{-1})$			BF(%)		
	RQDO ^a	RQDO ^b	SE ^c	RQDO ^a	RQDO ^b	SE ^c
$5s5p\ ^3P_0 - 5s5d\ ^3D_1$	0.696	0.593	0.665	57.9	57.9	57.17
$5s5p\ ^3P_1 -$	0.483	0.412	0.470	40.2	40.2	40.38
$5s5p\ ^3P_2 -$	0.023	0.020	0.028	1.9	1.9	2.38
$5s5p\ ^1P_1 -$	-	-	0.001	-	-	0.07
$5s5p\ ^3P_1 - 5s5d\ ^3D_2$	0.873	0.747	0.849	80.1	80.1	77.22
$5s5p\ ^3P_2 -$	0.218	0.186	0.249	19.9	19.9	22.70
$5s5p\ ^1P_1 -$	-	-	0.001	-	-	0.08
$5s5p\ ^3P_2 - 5s5d\ ^3D_3$	0.878	0.754	1.063	100.	100.	100.

^a Lavín and Martin, rel. quantum defect orbital [15].

^b Lavín and Martin, rel. quantum defect orbital with polarization [15].

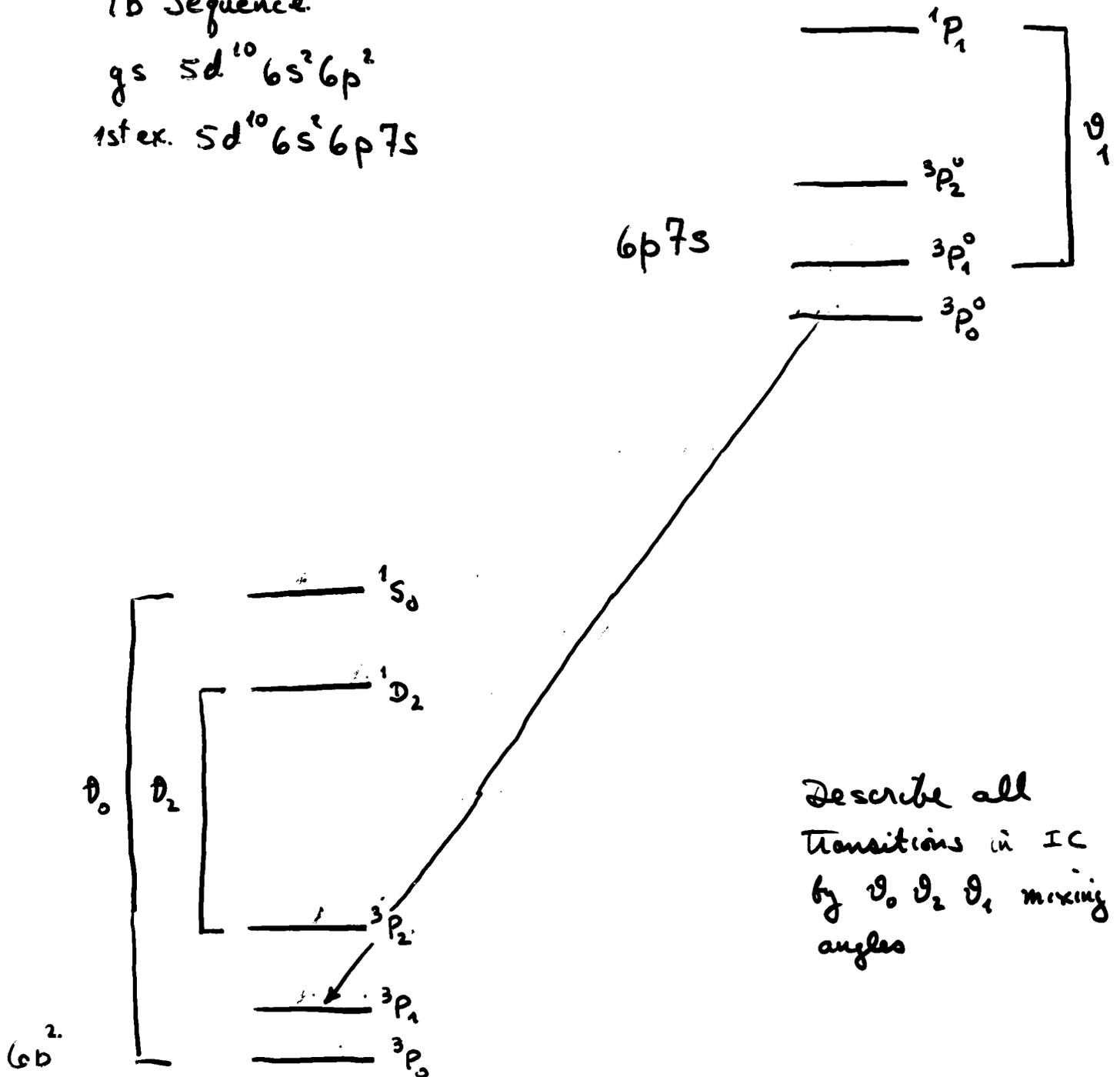
^c This work.

Example where transitions connect these two types of IC configurations

Pb Sequence.

gs $5d^{10}6s^26p^2$

1st ex. $5d^{10}6s^26p7s$



Describe all transitions in IC by θ_0 , θ_2 , θ_1 mixing angles

np^2

	3P_0	3P_1	3P_2	1D_2	1S_0
3P_0	$E_0 - 5F_2 - \zeta_1$	0	0	0	$-\sqrt{2} \zeta_0$
3P_1	0	$E_0 - 5F_2 - \zeta_1/2$	0	0	0
3P_2	0	0	$E_0 - 5F_2 + \zeta_1/2$	$\zeta_2/\sqrt{2}$	0
1D_2	0	0	$\zeta_2/\sqrt{2}$	$F_0 + F_2$	0
1S_0	$-\sqrt{2} \zeta_0$	0	0	0	$F_0 + 10F_2$

This is diagonalized by the transformation $J' = J$, where

$$J = \begin{pmatrix} \cos \theta_0 & 0 & 0 & 0 & \sin \theta_0 \\ 0 & 1 & \cos \theta_1 & \sin \theta_1 & 0 \\ 0 & 0 & -\sin \theta_1 & \cos \theta_1 & 0 \\ -\sin \theta_0 & 0 & 0 & 0 & \cos \theta_0 \end{pmatrix}$$

with the energy levels

$$^3P_0' = E_0 + 5F_2/2 - \zeta_1/2 - \Delta_0$$

$$^3P_1 = E_0 - 5F_2 - \zeta_1/2$$

$$^3P_2' = E_0 - 2F_2 + \zeta_1/4 - \Delta_2$$

$$^1D_2' = E_0 - 2F_2 + \zeta_1/4 + \Delta_2$$

$$^1S_0' = E_0 + 5F_2/2 - \zeta_1/2 + \Delta_2$$

$$\begin{cases} \Delta_0 = \sqrt{(15F_2/2 + \zeta_1/2)^2 + 2\zeta_0^2} \\ \Delta_2 = \sqrt{(3F_2 - \zeta_1/4)^2 + \zeta_2^2/2} \end{cases}$$

where the angles are

$$\cot 2\theta_0 = \frac{15F_2 + \zeta_1}{\sqrt{8} \zeta_0}$$

$$\cot 2\theta_2 = \frac{12F_2 - \zeta_1}{\sqrt{8} \zeta_2}$$

levels of the ground configuration p^2 can be deduced from this formalism using the LS coupling angular transition matrices [12, 13]. The nonvanishing values are

$$\langle {}^3P_0^o | r | {}^3P_1 \rangle = \langle {}^1P_1^o | r | {}^1S_0 \rangle = -\langle {}^3P_1^o | r | {}^3P_0 \rangle = -\sqrt{20} \quad (10)$$

$$2\langle {}^3P_2^o | r | {}^3P_1 \rangle = -2\langle {}^3P_1^o | r | {}^3P_2 \rangle = \langle {}^1P_1^o | r | {}^1D_2 \rangle = 10 \quad (11)$$

$$\sqrt{5}\langle {}^3P_1^o | r | {}^3P_1 \rangle = \langle {}^3P_2^o | r | {}^3P_2 \rangle = \sqrt{75}. \quad (12)$$

These equations yield, for the upper level ${}^3P_1^o$

$$\langle {}^3P_0' | \mathbf{r} | {}^3P_1^o \rangle = -\sqrt{20} \cos(\theta_1 + \theta_0) \langle p^2 | r | sp \rangle \quad (13)$$

$$\langle {}^3P_1' | \mathbf{r} | {}^3P_1^o \rangle = \sqrt{15} \cos \theta_1 \langle p^2 | r | sp \rangle \quad (14)$$

$$\langle {}^3P_2' | \mathbf{r} | {}^3P_1^o \rangle = 5(2 \sin \theta_1 \sin \theta_2 + \cos \theta_1 \cos \theta_2) \langle p^2 | r | sp \rangle \quad (15)$$

$$\langle {}^1D_2' | \mathbf{r} | {}^3P_1^o \rangle = -5(2 \sin \theta_1 \cos \theta_2 - \cos \theta_1 \sin \theta_2) \langle p^2 | r | sp \rangle \quad (16)$$

$$\langle {}^1S_0' | \mathbf{r} | {}^3P_1^o \rangle = -\sqrt{20} \sin(\theta_1 + \theta_0) \langle p^2 | r | sp \rangle, \quad (17)$$

for the upper level ${}^3P_2^o$

$$\langle {}^3P_1' | \mathbf{r} | {}^3P_2^o \rangle = 5 \langle p^2 | r | sp \rangle \quad (18)$$

$$\langle {}^3P_2' | \mathbf{r} | {}^3P_2^o \rangle = \sqrt{15} \cos \theta_2 \langle p^2 | r | sp \rangle \quad (19)$$

$$\langle {}^1D_2' | \mathbf{r} | {}^3P_2^o \rangle = \sqrt{15} \sin \theta_2 \langle p^2 | r | sp \rangle, \quad (20)$$

and for the upper level ${}^1P_1^o$

$$\langle {}^3P_0' | \mathbf{r} | {}^1P_1^o \rangle = -\sqrt{20} \sin(\theta_1 + \theta_0) \langle p^2 | r | sp \rangle \quad (21)$$

$$\langle {}^3P_1' | \mathbf{r} | {}^1P_1^o \rangle = \sqrt{15} \sin \theta_1 \langle p^2 | r | sp \rangle \quad (22)$$

$$\langle {}^3P_2' | \mathbf{r} | {}^1P_1^o \rangle = 5(2 \cos \theta_1 \sin \theta_2 - \sin \theta_1 \cos \theta_2) \langle p^2 | r | sp \rangle \quad (23)$$

$$\langle {}^1D_2' | \mathbf{r} | {}^1P_1^o \rangle = -5(2 \cos \theta_1 \cos \theta_2 + \sin \theta_1 \sin \theta_2) \langle p^2 | r | sp \rangle \quad (24)$$

$$\langle {}^1S_0' | \mathbf{r} | {}^1P_1^o \rangle = -\sqrt{20} \cos(\theta_1 + \theta_0) \langle p^2 | r | sp \rangle. \quad (25)$$

It should be noted that in a fully relativistic Dirac treatment the corresponding expressions will involve two separate jj coupled radial transition matrices, and reduce to equations (13-25) only if these two radial matrices are equal. Theoretical studies of these relativistic corrections have been presented elsewhere [14].

For pure sp and p^2 configurations the energy levels (and thereby the mixing angles) are specified [10] by three parameters (F_0, G_1, ζ_p for sp and F_0, F_2, ζ_{pp} for p^2 , in the notation of Ref.[13]). Since the sp and p^2 configurations contain four and five levels respectively, the specification of these three parameters is overdetermined. Here this was treated by using the average energies ε_J of the $J=0,1,2$ levels to make an exactly determined parametrization, computing the singlet-triplet splittings from this parametrization, and then using the deviations as a measure of the validity of the

Table 1. Comparison of semiempirical and measured branching fractions (in %) for Si I and Ge I. SE denotes the semiempirical estimates of [3]. Expt denotes experimental measurements as cited, with parentheses indicating quoted uncertainties in the last figure.

Transition	Si I			Ge I		
	SE	Expt [6]	Expt [7]	SE	Expt [8]	Expt [9]
$^3P'_0 \leftarrow ^3P'_1$	33.3	33.3(17)	33.3(3)	31.2	32.5(16)	32.9
$^3P_1 \leftarrow$	24.7	24.7(13)	24.7(4)	21.2	22.1(11)	20.3
$^3P'_2 \leftarrow$	41.1	40.6(21)	40.7(4)	38.3	37.1(19)	36.1
$^1D'_2 \leftarrow$	0.88	1.20(11)	1.2(1)	8.8	8.1(8)	10.3
$^1S'_0 \leftarrow$	0.06	<0.20(6)	<0.20(6)	0.52	0.23(2)	0.38
$^3P_1 \leftarrow ^3P^0_2$	25.2	24.6(13)	24.6(3)	26.4	27.2(14)	31.0
$^3P'_2 \leftarrow$	74.8	75.4(36)	75.4(3)	73.1	72.1(14)	67.8
$^1D'_2 \leftarrow$	0.020	0.027(4)	0.027(4)	0.53	0.72(7)	1.3
$^3P'_0 \leftarrow ^1P^0_1$	0.24	0.34(3)	0.30(2)	2.9	4.6(5)	4.5
$^3P_1 \leftarrow$	0.25	0.27(3)	0.20(2)	3.3	3.6(4)	3.6
$^3P'_2 \leftarrow$	0.15	0.25(3)	0.20(2)	1.0	1.68(17)	1.7
$^1D'_2 \leftarrow$	92.0	93.4(47)	93.4(2)	86.2	86.1(14)	83.2
$^1S'_0 \leftarrow$	7.4	5.7(3)	5.70(12)	6.6	4.0(4)	7.0

Table 4. Transition wavelengths and semiempirical branching fractions (BF, in %) for ions in the Si sequence.

Transition	P II		S III		Cl IV		Ar V	
	λ (Å)	BF	λ (Å)	BF	λ (Å)	BF	λ (Å)	BF
$^3P'_0 \leftarrow ^3P'_1$	1115.82	33.1	681.49	32.8	464.28	32.4	337.58	31.9
$^3P_1 \leftarrow$	1155.01	24.5	682.88	24.2	465.34	23.8	338.45	23.2
$^3P'_2 \leftarrow$	1158.82	41.0	685.38	40.8	467.19	40.7	339.91	40.5
$^1D'_2 \leftarrow$	1284.33	1.3	784.47	1.9	495.98	2.7	357.24	3.8
$^1S'_0 \leftarrow$	1534.49	0.12	836.28	0.22	546.92	0.32	387.12	0.49
$^3P_1 \leftarrow ^3P'_2$	1149.96	25.2	680.97	25.2	463.01	25.2	336.57	25.3
$^3P'_2 \leftarrow$	1153.73	74.8	683.46	74.7	464.84	74.6	338.01	74.4
$^1D'_2 \leftarrow$	1278.08	$<10^{-3}$	736.25	0.11	493.34	0.21	355.14	0.37
$^3P'_0 \leftarrow ^1P'_1$	1124.95	0.22	673.86	0.27	455.68	0.34	331.91	0.47
$^3P_1 \leftarrow$	1127.04	0.25	675.22	0.34	456.70	0.46	332.75	0.65
$^3P'_2 \leftarrow$	1130.66	0.11	677.66	$<10^{-3}$	458.48	$<10^{-3}$	334.16	$<10^{-3}$
$^1D'_2 \leftarrow$	1249.83	88.8	729.52	87.2	486.18	86.1	350.89	85.2
$^1S'_0 \leftarrow$	1485.50	10.6	824.82	12.1	535.03	13.0	379.68	13.6

Table 5. Transition wavelengths and semiempirical branching fractions (in %) for ions in the Ge sequence.

Transition	As II		Se III		Br IV	
	λ (Å)	BF	λ (Å)	BF	λ (Å)	BF
$^3P'_0 \leftarrow ^3P'_1$	1263.77	29.8	788.76	28.1	549.93	28.8
$^3P_1 \leftarrow$	1280.99	20.0	799.74	18.4	557.98	18.6
$^3P'_2 \leftarrow$	1305.70	38.7	814.04	38.6	567.45	40.8
$^1D'_2 \leftarrow$	1448.59	10.5	879.13	13.4	603.68	10.5
$^1S'_0 \leftarrow$	1768.98	0.99	1017.30	1.5	676.48	1.3
$^3P_1 \leftarrow ^3P'_2$	1243.08	26.2	777.30	26.1	545.44	26.1
$^3P'_2 \leftarrow$	1266.34	72.7	790.80	72.0	554.48	70.8
$^1D'_2 \leftarrow$	1400.30	1.1	852.09	1.87	589.03	3.1
$^3P'_0 \leftarrow ^1P'_1$	1207.45	2.34	759.56	2.70	534.47	1.9
$^3P_1 \leftarrow$	1223.16	2.95	769.74	3.50	542.07	3.0
$^3P'_2 \leftarrow$	1245.67	0.41	782.97	0.32	551.00	$<10^{-5}$
$^1D'_2 \leftarrow$	1375.07	84.4	843.00	82.2	585.08	82.6
$^1S'_0 \leftarrow$	1660.56	9.90	968.24	11.3	663.23	12.6

Table III. Wavelengths and semiempirical, theoretical and experimental branching fractions for $5s^25p^2-5s^25p6s$ transitions in Sn I and Sb II.

Transition	Sn I						Sb II		
	$\lambda(\text{\AA})^a$	BF(%)					$\lambda(\text{\AA})^a$	BF(%)	
		SE ^b	B ^c	L ^d	CB ^e	M ^f		SE ^b	B ^c
$^3P_0 \leftarrow ^3P_1^o$	2863.32	32.3	28.9	27	37	40	1438.11	30.2	38.4
$^3P_1 \leftarrow$	3009.13	17.5	20.9	17	27	28	1504.19	16.1	23.8
$^3P_2 \leftarrow$	3175.03	39.7	41.8	39	22	22	1565.50	43.4	31.9
$^1D_2 \leftarrow$	3801.01	10.0	8.2	17	14	11	1762.24	9.5	5.7
$^1S_0 \leftarrow$	5631.71	0.5	0.2	-	0.3	0.2	2190.85	0.8	0.2
$^3P_1 \leftarrow ^3P_2^o$	2706.50	28.3	25.3	22	30	32	1384.66	27.8	14.8
$^3P_2 \leftarrow$	2839.98	68.5	72.1	71	64	63	1436.45	65.5	48.4
$^1D_2 \leftarrow$	3330.61	3.2	2.6	7	6	5	1600.39	6.6	36.8
$^3P_0 \leftarrow ^1P_1^o$	2546.55	4.2	3.6	8	21	25	1317.54	3.0	1.6
$^3P_1 \leftarrow$	2661.24	6.8	3.6	4	13	14	1372.79	6.0	1.8
$^3P_2 \leftarrow$	2790.18	0.01	0.8	-	-	-	1423.68	0.6	0.5
$^1D_2 \leftarrow$	3262.33	82.2	84.3	88	60	57	1584.56	80.2	84.7
$^1S_0 \leftarrow$	4624.75	6.8	7.7	-	6	4	1923.32	10.2	11.5

^a Air wavelengths for $\lambda > 2000 \text{ \AA}$.

^b Semiempirical, this work.

^c Bieron *et al.* [18], theoretical (MCDM-EAL with Babushkin gauge).

^d Lotrian *et al.* [19], arc emission.

^e Corliss and Bozman [20], arc emission.

^f Meggers *et al.* [21], arc emission.

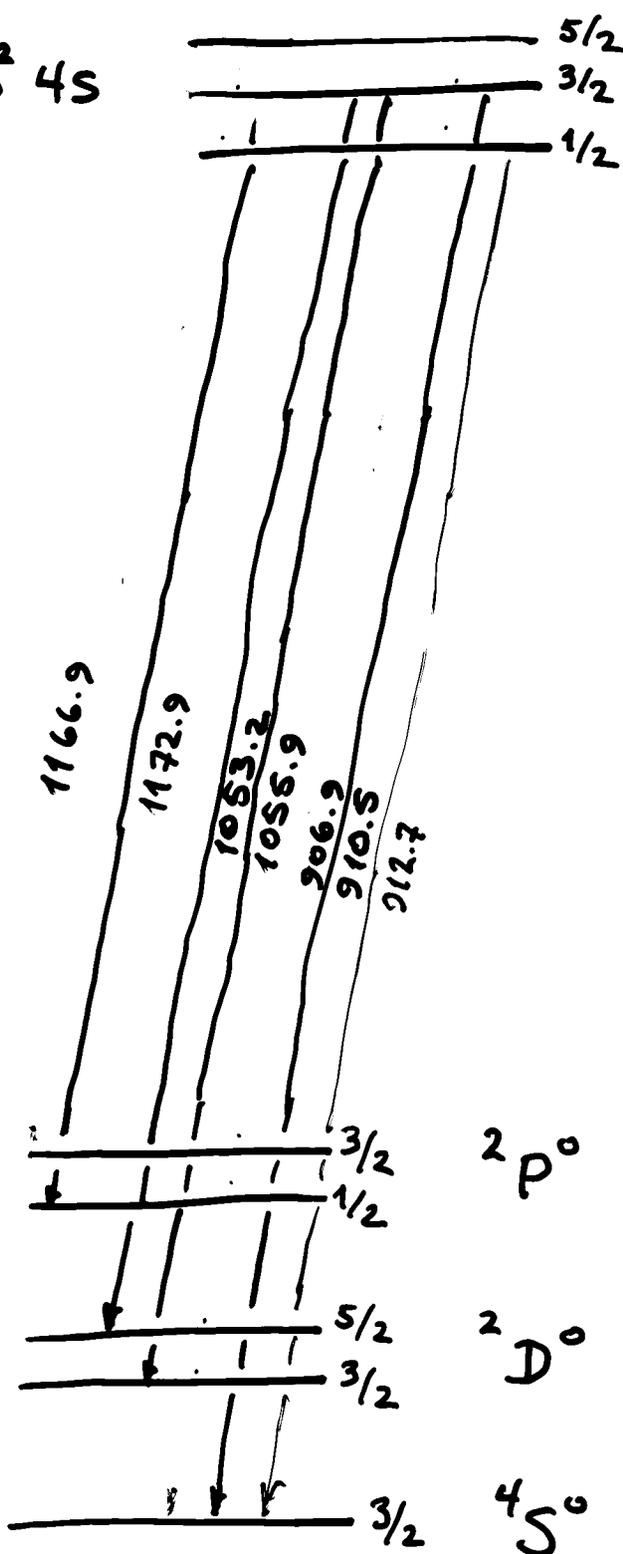
Table IV. Wavelengths and semiempirical branching fractions for $5s^25p^2-5s^25p6s$ transitions in multiply charged ions of the Sn sequence.

Transition	Te III		I IV		Xe V		Cs VI	
	$\lambda(\text{\AA})$	BF(%)	$\lambda(\text{\AA})$	BF(%)	$\lambda(\text{\AA})$	BF(%)	$\lambda(\text{\AA})$	BF(%)
$^3P_0 \leftarrow ^3P_1^o$	928.30	30.1	666.29	30.0	512.82	31.1	410.31	31.0
$^3P_1 \leftarrow$	971.19	15.6	698.04	15.3	538.56	16.1	431.89	15.6
$^3P_2 \leftarrow$	1004.45	45.1	718.89	49.0	552.94	50.0	442.30	51.2
$^1D_2 \leftarrow$	1106.63	8.5	783.98	5.0	600.38	2.4	479.25	1.8
$^1S_0 \leftarrow$	1310.56	0.7	885.67	0.7	664.81	0.4	522.72	0.3
$^3P_1 \leftarrow ^3P_2^o$	913.59	27.6	649.31	27.6	500.55	27.6	401.97	27.7
$^3P_2 \leftarrow$	942.97	63.6	667.31	57.8	512.95	54.3	410.98	51.1
$^1D_2 \leftarrow$	1032.46	8.9	723.04	14.6	553.53	18.1	442.09	21.2
$^3P_0 \leftarrow ^1P_1^o$	866.40	2.1	619.65	1.5	469.19	0.7	378.86	0.6
$^3P_1 \leftarrow$	903.64	5.6	647.02	5.2	490.64	4.3	396.75	4.4
$^3P_2 \leftarrow$	932.37	1.4	664.89	5.2	502.55	9.4	405.52	11.9
$^1D_2 \leftarrow$	1019.77	79.2	720.20	74.8	541.43	71.6	456.37	68.6
$^1S_0 \leftarrow$	1190.47	11.7	805.12	13.3	593.28	14.1	472.11	14.6

S II

$3s^2 3p^2 4s$

4p



Te III	
903.6	5.6%
913.6	27.6%
1032.5	8.9%
1190.5	11.7%

$3s^2 3p^3$

$2p^o$

$2D^o$

$4S^o$

Dirac Formulation

$$\langle {}^3P_1 | \mathbf{r} | {}^3P_0^\circ \rangle = \sqrt{20}R_{31} \quad (1)$$

$$\begin{aligned} \langle {}^3P_0 | \mathbf{r} | {}^3P_1^\circ \rangle = & -\frac{\sqrt{20}}{3} [(R_{31} + 2R_{11}) \cos \theta_0 \cos \theta_1 - (2R_{31} + R_{11}) \sin \theta_0 \sin \theta_1 \\ & + \sqrt{2}(R_{31} - R_{11}) \sin(\theta_0 - \theta_1)] \end{aligned} \quad (2)$$

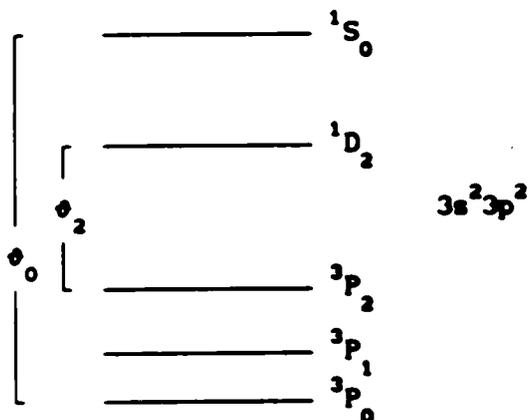
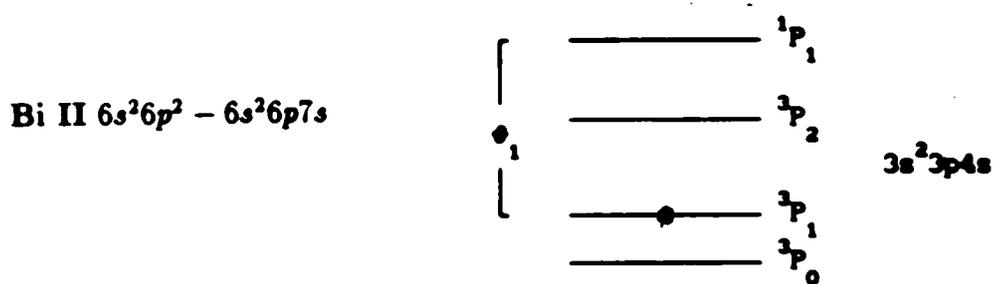
$$\langle {}^3P_1 | \mathbf{r} | {}^3P_1^\circ \rangle = \frac{\sqrt{15}}{3} [(2R_{31} + R_{11}) \cos \theta_1 + \sqrt{2}(R_{31} - R_{11}) \sin \theta_1] \quad (3)$$

$$\begin{aligned} \langle {}^3P_2 | \mathbf{r} | {}^3P_1^\circ \rangle = & \frac{5}{3} [(4R_{31} + 2R_{11}) \sin \theta_1 \sin \theta_2 + (4R_{31} - R_{11}) \cos \theta_1 \cos \theta_2 \\ & - \sqrt{2}(R_{31} - R_{11}) \sin(\theta_1 - \theta_2)] \end{aligned} \quad (4)$$

$$\begin{aligned} \langle {}^1D_2 | \mathbf{r} | {}^3P_1^\circ \rangle = & -\frac{5}{3} [(4R_{31} + 2R_{11}) \sin \theta_1 \cos \theta_2 - (4R_{31} - R_{11}) \cos \theta_1 \sin \theta_2 \\ & + \sqrt{2}(R_{31} - R_{11}) \cos(\theta_1 - \theta_2)] \end{aligned} \quad (5)$$

$$\begin{aligned} \langle {}^1S_0 | \mathbf{r} | {}^3P_1^\circ \rangle = & -\frac{\sqrt{20}}{3} [(R_{31} + 2R_{11}) \sin \theta_0 \cos \theta_1 + (2R_{31} + R_{11}) \cos \theta_0 \sin \theta_1 \\ & - \sqrt{2}(R_{31} - R_{11}) \cos(\theta_0 - \theta_1)] \end{aligned} \quad (6)$$

Intermediate Coupling Specification of Branching Fractions



$$\langle {}^3P_0 | r | {}^3P_1^{\circ} \rangle = -\sqrt{20} \cos(\theta_1 + \theta_0) \langle p^2 | r | sp \rangle$$

$$\langle {}^3P_1 | r | {}^3P_1^{\circ} \rangle = \sqrt{15} \cos \theta_1 \langle p^2 | r | sp \rangle$$

$$\langle {}^3P_2 | r | {}^3P_1^{\circ} \rangle = 5(2 \sin \theta_1 \sin \theta_2 + \cos \theta_1 \cos \theta_2) \langle p^2 | r | sp \rangle$$

$$\langle {}^1D_2 | r | {}^3P_1^{\circ} \rangle = -5(2 \sin \theta_1 \cos \theta_2 - \cos \theta_1 \sin \theta_2) \langle p^2 | r | sp \rangle$$

$$\langle {}^1S_0 | r | {}^3P_1^{\circ} \rangle = -\sqrt{20} \sin(\theta_1 + \theta_0) \langle p^2 | r | sp \rangle .$$

Table 2: Pb I branching fractions and transition probability rates for the $^3P_1^o$ upper level in the $6s^26p^2-6s^26p7s$ multiplet.

Transition	$\lambda(\text{\AA})^a$	BF(N) ^b	BF(R) ^c	BF(M) ^d	$A(ns^{-1})^e$
$^3P_0 - ^3P_1^o$	2833.89	0.489	0.310	0.324	0.0529
$^3P_1 -$	3640.61	0.128	0.166	0.188	0.0284
$^3P_2 -$	4058.95	0.381	0.520	0.500	0.0889
$^1D_2 -$	7230.96	0.0029	0.0040	0.0005	0.00068
$^1S_0 -$	17181	7×10^{-5}	3×10^{-5}	-	6×10^{-6}

^a Vacuum wavelengths.

^b Nonrelativistic, $R_{13}/R_{11}=1$.

^c Relativistic, $R_{13}/R_{11}=1.4590$.

^d Measured, Ref.[22].

^e Relativistic, using BF(R) and $\tau=5.84$ ns.

Table 3: Bi II branching fractions and transition probability rates for the $^3P_1^o$ upper level in the $6s^26p^2-6s^26p7s$ multiplet.

Transition	$\lambda(\text{\AA})^a$	BF(N) ^b	BF(R) ^c	$A(ns^{-1})^d$
$^3P_0 - ^3P_1^o$	1436.83 ^e	0.43	0.25	0.20
$^3P_1 -$	1777.11 ^e	0.12	0.16	0.13
$^3P_2 -$	1902.31 ^f	0.44	0.59	0.47
$^1D_2 -$	2804.2 ^e	0.004	0.005	0.004
$^1S_0 -$	3933.3 ^e	0.0002	0.0009	0.0002

^a Vacuum wavelengths.

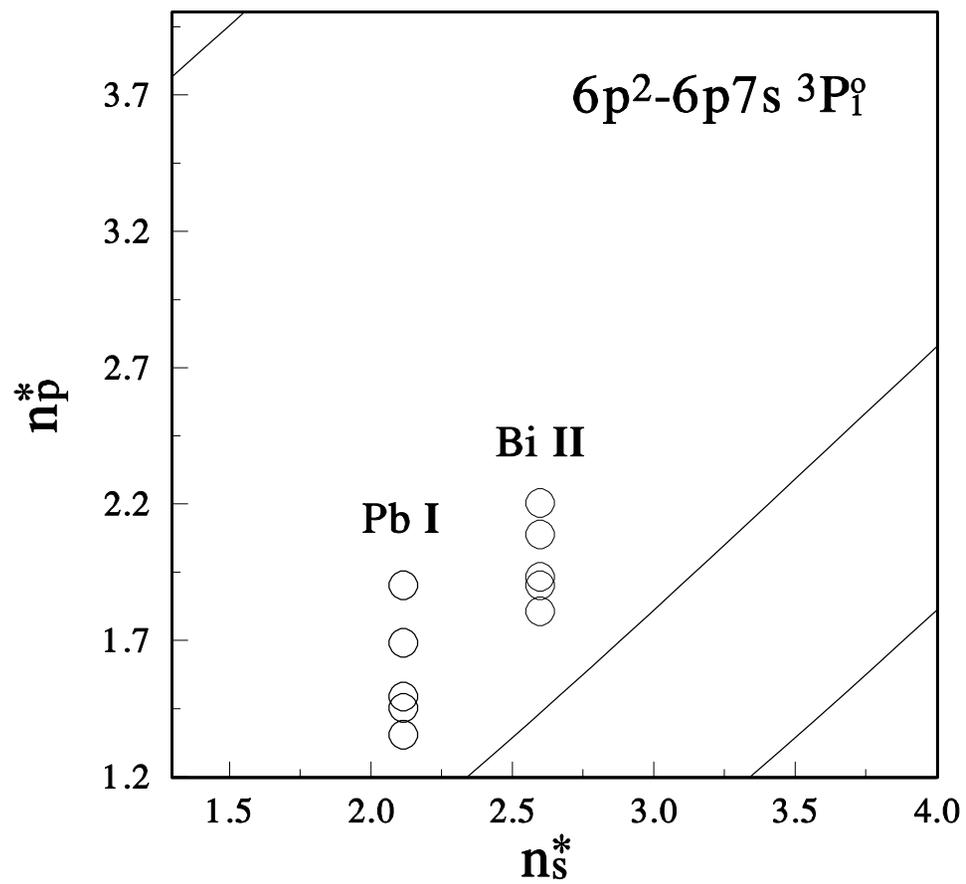
^b Nonrelativistic, $R_{13}/R_{11}=1$.

^c Relativistic, $R_{13}/R_{11}=1.4224$.

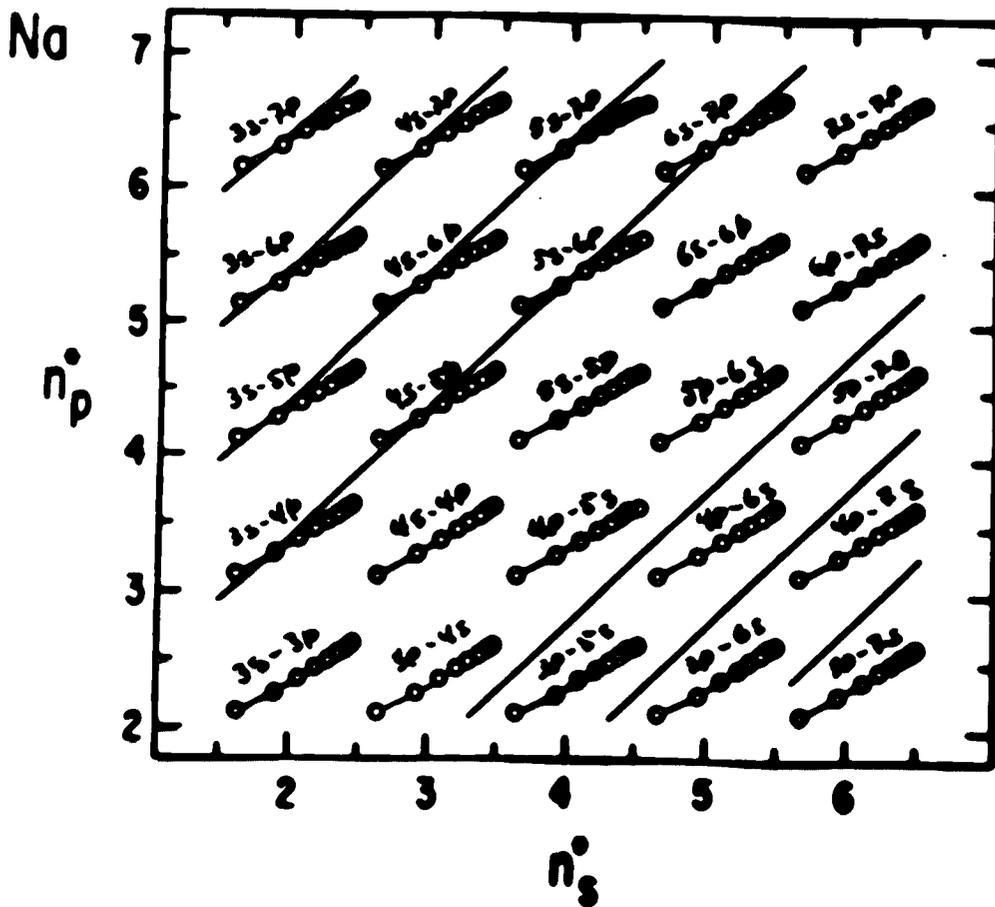
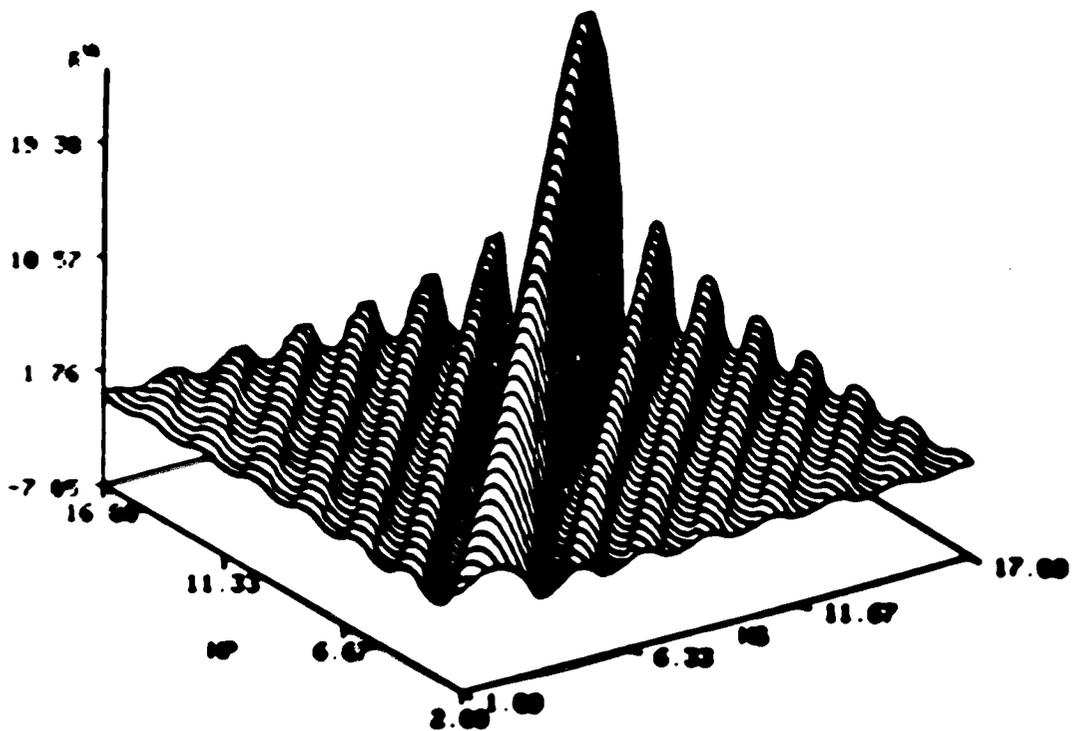
^d Relativistic, using BF(R) and $\tau=1.56$ ns.

^e Reader and Corliss, Ref.[26].

^f Wahlgren *et al.*, Ref.[27].



DIPOLE $ns-np$



DIPOLE $ns-np$

